



# *The Application of R-Matrix Analysis to Experimental Data*

## 1 - Resonance Properties

Alex Murphy & Alison Laird



**David Mountford**



# Contents

- The R-Matrix Formalism
- Context: The  $^{18}\text{F}(\text{p},\alpha)^{15}\text{O}$  Reaction
- (Doing an...) R-Matrix analysis
- Comparison of R-Matrix Codes
- Conclusions



# R-Matrix

The  
experimentalist's  
hope



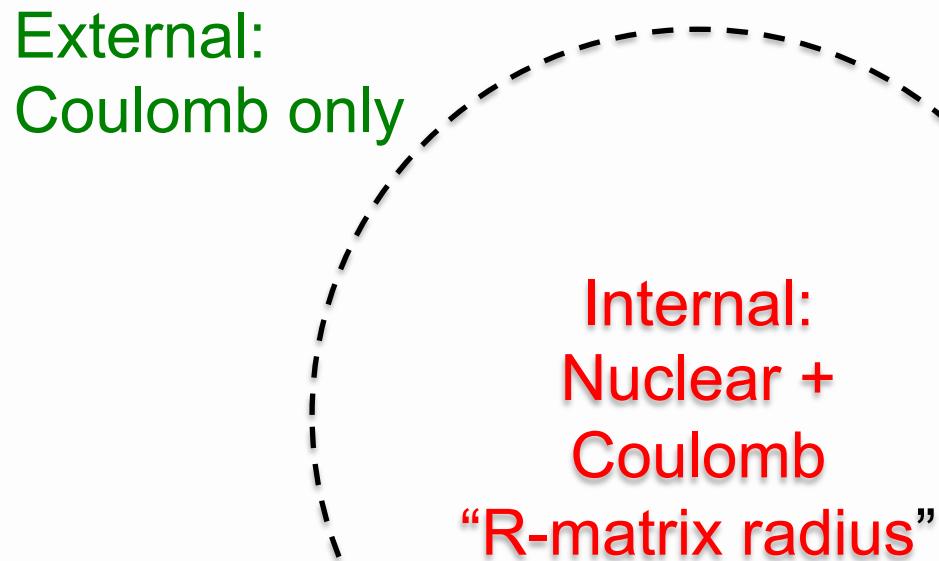
...but it's never  
quite that easy...



**Some Theory...**

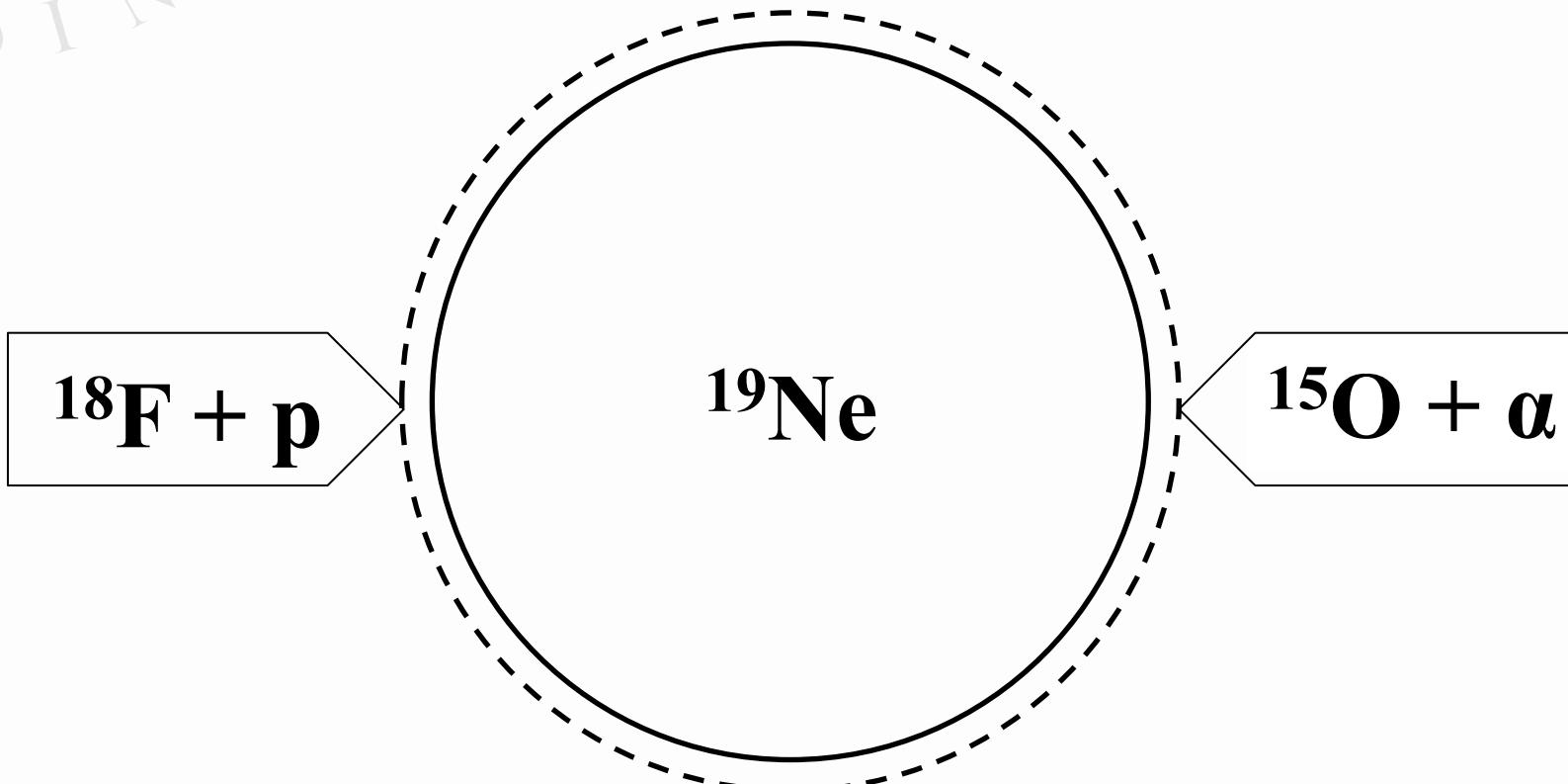
# The R-Matrix Formalism

- “Bible”: Lane and Thomas, Rev. Mod. Phys. 30, 257 (1958)[1]
- Basics: split system into internal and external regions:



# The R-Matrix Formalism

- R-Matrix radius  $\sim$  nuclear radius, slightly larger
- Various incoming “channels”



# The R-Matrix Formalism

- Internal Region:
  - Wavefunction not well understood, cannot obtain cross section...
    - Radial and spherical contributions
    - Set of orthonormal basis functions

$$\Psi = \sum_c \psi_c \phi_c = \sum_\lambda C_\lambda \chi_\lambda$$

- Calculating coefficients and solving Schrodinger equation:

$$\phi_c \left( \frac{\hbar^2}{2\mu_c r_c} \right)^{1/2} = \sum_\lambda (E_\lambda - E)^{-1} \sum_{c'} \left( \frac{\hbar^2}{2\mu_{c'} r_{c'}} \right)^{1/2} \gamma_{\lambda c} \gamma_{\lambda c'} [\rho_{c'} \phi_{c'} - \phi_{c'} b_{c'}]$$

- Purely radial
- All spherical dependence absorbed by  $\gamma$
- Extract “R-matrix”:

$$R_{cc'} = \sum_\lambda \frac{\gamma_{\lambda c} \gamma_{\lambda c'}}{E_\lambda - E}$$

# The R-Matrix Formalism

- External Region:
  - Purely Coulomb potential
  - Radial wavefunction is well understood combination of well known Coulomb functions:

$$\phi_c = \frac{1}{\sqrt{v_c}} \left( A_c I_c - \sum_{c'} U_{cc'} A_{c'} O_c \right)$$

$$I_c = (G_c - iF_c) e^{i\omega_c}$$

$$O_c = (G_c + iF_c) e^{-i\omega_c}$$

- $U_{cc'}$  = scattering matrix
- Substitution into internal wavefunction gives scattering matrix in terms of R-matrix:

$$U = \rho^{1/2} O^{-1} (1 - RL)^{-1} (1 - RL^*) I \rho^{-1/2}$$

# The R-Matrix Formalism

- To cross section:
  - One  $U$  for every  $J^\pi$  group

$$\frac{d\sigma}{d\Omega} \propto |U_J U_{J'}^*|$$

- $U \sim e^{2i\delta}$ ,  $\delta$  = total phase shift:

$$\frac{d\sigma}{d\Omega} \propto \cos(2(\delta_J - \delta_{J'}))$$

- $\delta$  = phase of  $U$
- Cross section calculated either directly from  $U$  or from phase shifts

# R-Matrix Implementation

- Simplified multi channel code (P. Descouvemont):

$$\frac{d\sigma}{d\Omega} \propto \cos(2(\delta_J - \delta_{J'}))$$

*'Still Living the DREAM'*

A	B	C	D	E	F	G	H	I	J	K	L	M	N	O	Q
1 A1=	18	A2=	1	Z1=	9	Z2=	1	a1=	15	a2=	4	z1=	8	z2=	
2 I1=	1	I2=	1				i1=	1	I2=	1	0				
3 rmax=	5	lmax=	4				Q=	-2.8818							
4 hbar^2/2m=	20.736														
5															
6 J=															
7 ell=															
8 I=															
9 nres=															
10 Er=															
11 I' or y2=															
12 sign=															
13	-0.187271	0.7542637	1.1547151	1.245944	1.188422	0.967209	-0.652918	1.534383							
14	0.8824691	0.0234101	0.0708146	0.242541	0.319927	0.42501	0.71013	0.051694							
15	0.1455901	0.0114092	0.0067981	-0.140538	0.415444	0.009837	0.461131	0.039068							
16	4 partial wave(s)														
17	8 resonance(s)														
18															
19	DataSheet	elastic	inelastic												
20	Angle	175.63365	175.63365												
21	Type	0	1												
22	nbre points	54	47												
23	delta_E	0.003/0.005	0.007/0.013												
24	selection	1	1												
25	famul	1	1												
26	chi2 part.	1.9872439	1.2270268												
27															
28	Cells														
29	nval														
30	initial														
31	final														
32															
33		1.6334795		0.665	0.015	0.024	0.759	0.0016	0.0024	1.16	0.0023	0.0019	1.335	0.065	0.026
34	Iterations	1	errors												1.455
35	Angle	Angle cm	Ecm (MeV)	Nucl (mb)	Coul (mb)	Nucl/Coul	Inelastic								
36	175.6337	175.63365	0.585	287.65044	307.6378	0.93503	287.6504								
37	175.6337	175.63365	0.61	250.56181	282.9383	0.88557	250.5618								
38	175.6337	175.63365	0.635	210.04863	261.0983	0.804481	210.0486								
39	175.6337	175.63365	0.66	236.56629	241.6927	0.97879	236.5663								
40	175.6337	175.63365	0.685	316.90475	224.3728	1.412402	316.9047								

Normal View   Ready   HELP   Input   elastic   inelastic   Main   Sheet1   +   Sum=0



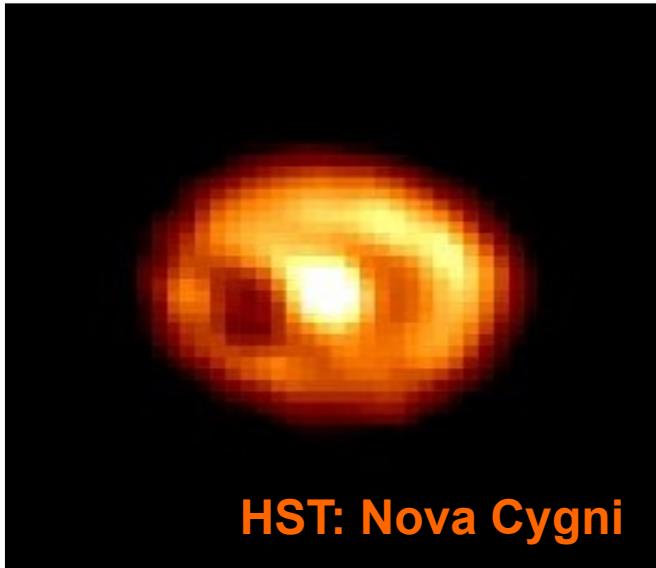
# Experimental Work



## Motivation

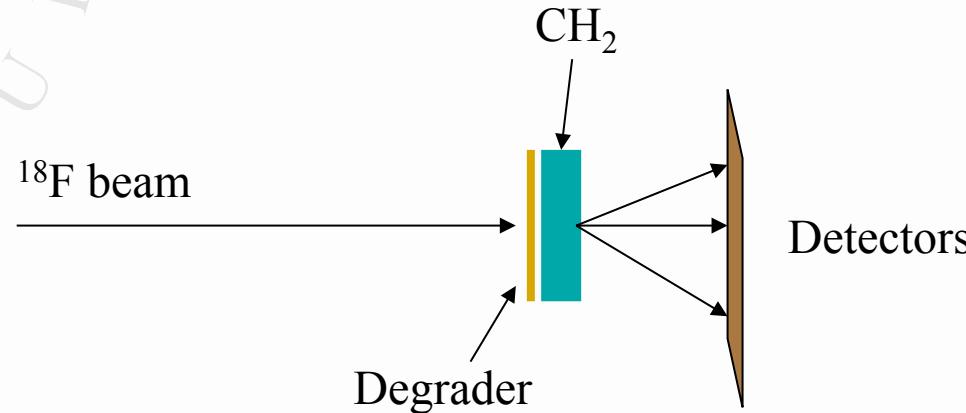
- $^{18}\text{F}$  is the dominant gamma-ray emitter in novae
- Abundance strongly dependent on rate of this reaction
- Reaction proceeds through poorly known resonances in  $^{19}\text{Ne}$
- Theoretical expectation of additional states

**Need to find and characterise states in  $^{19}\text{Ne}$**





# Generic Set Up

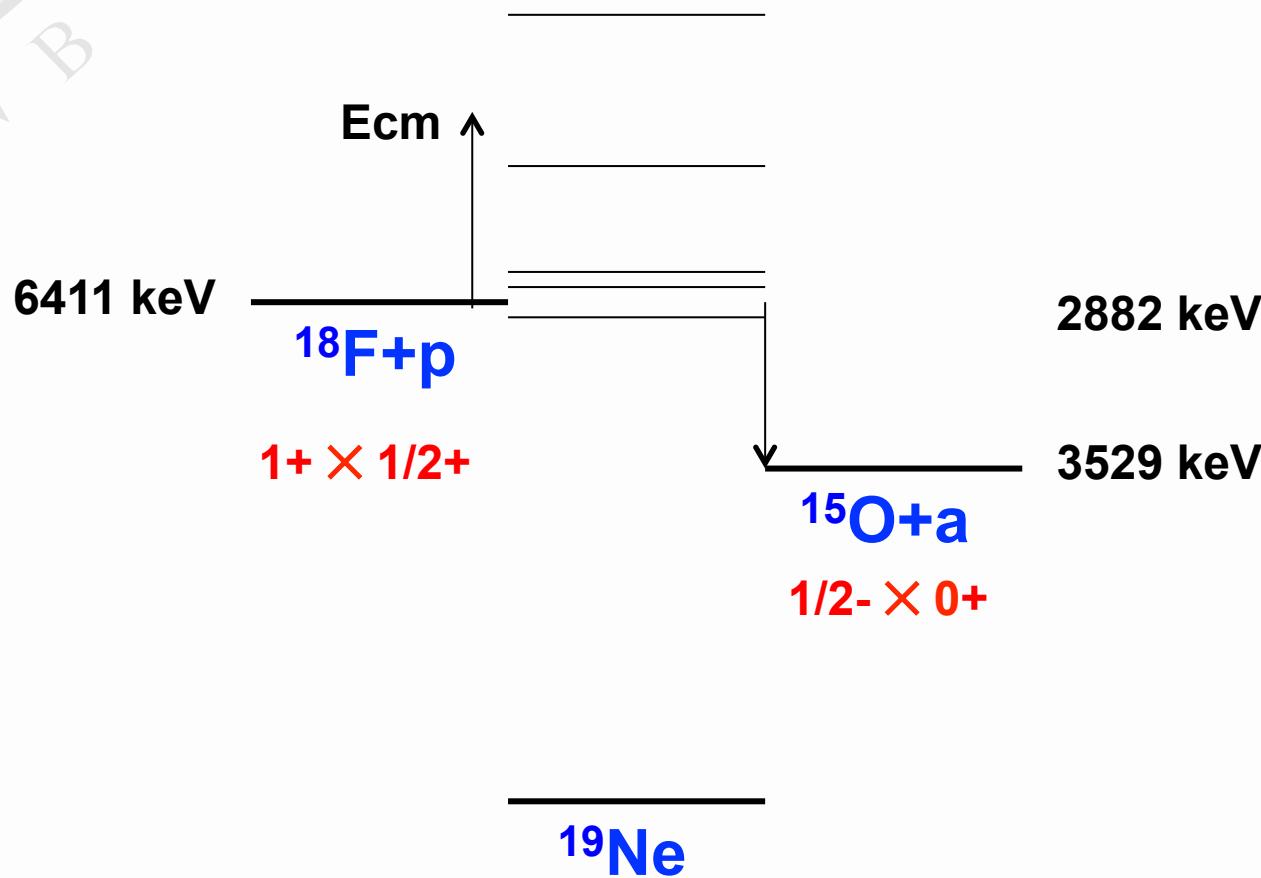


- Pure, intense  $^{18}\text{F}$  beam
- ‘Thick’ CH<sub>2</sub> target
- $^{18}\text{F}$  stopped in target
- Adjust beam energy & target thickness for desired  $E_{cm}$  coverage
- Protons and alpha particles detected downstream in DSSDs

**Simultaneous excitations functions of  
 $^{18}\text{F}(\text{p},\text{p})^{18}\text{F}$  &  $^{18}\text{F}(\text{p},\alpha)^{15}\text{O}$**

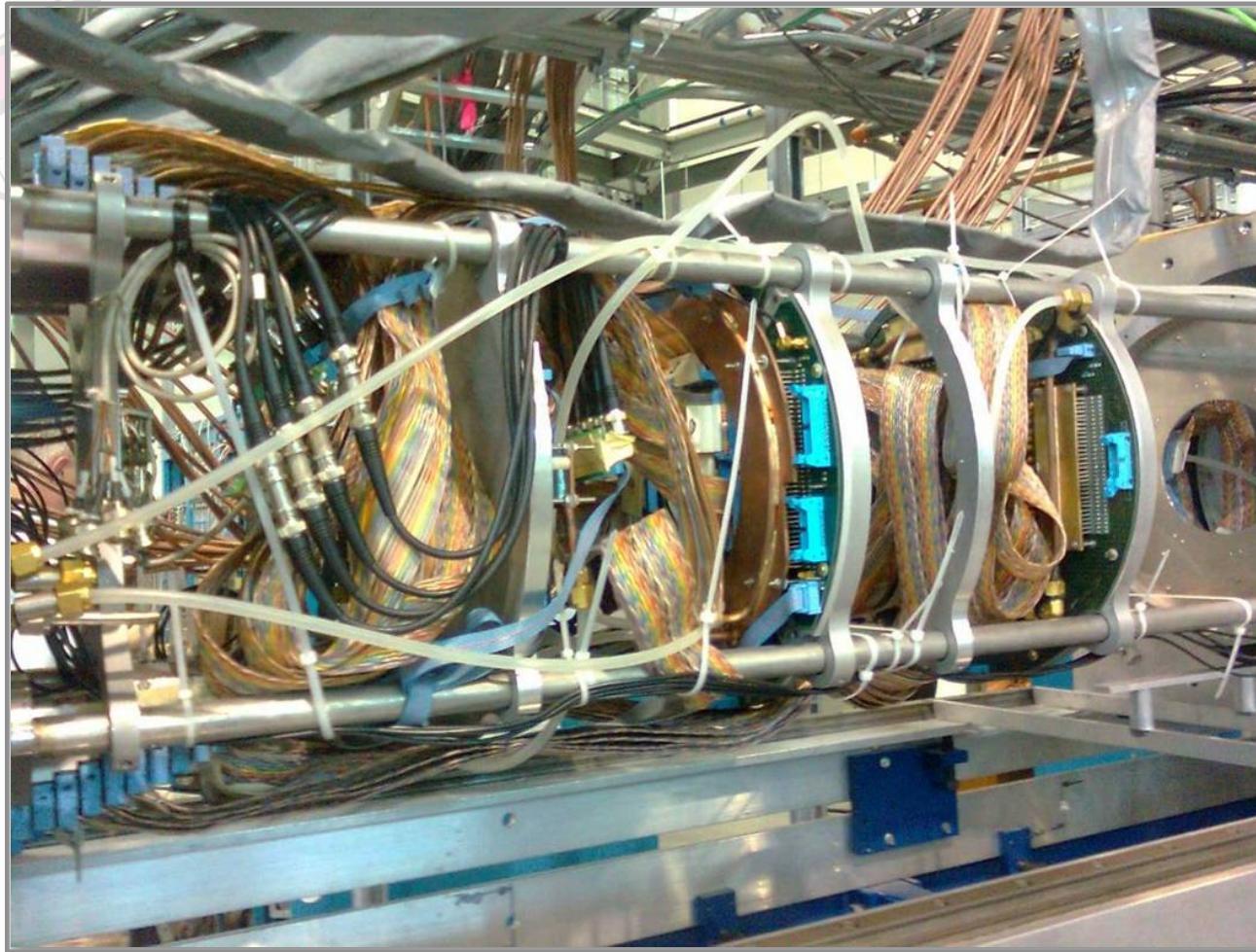


# Physics

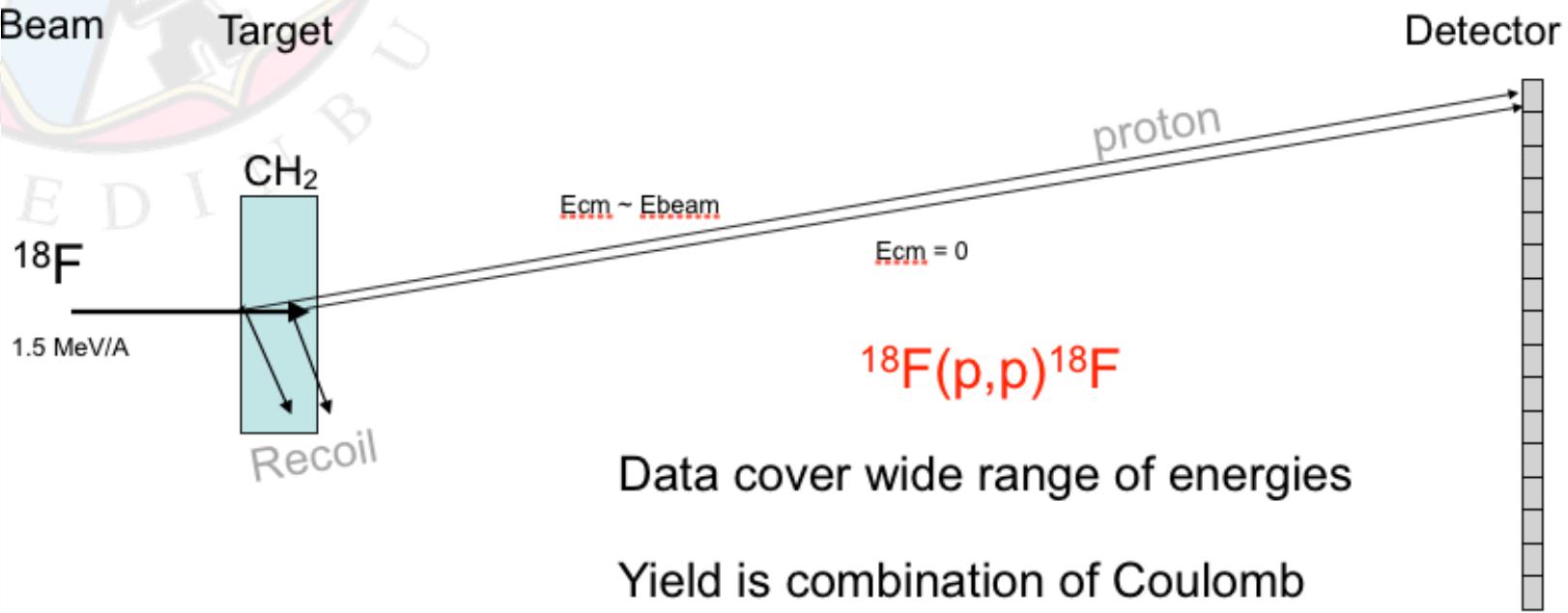




# Actual Set Up



# Thick Target Technique

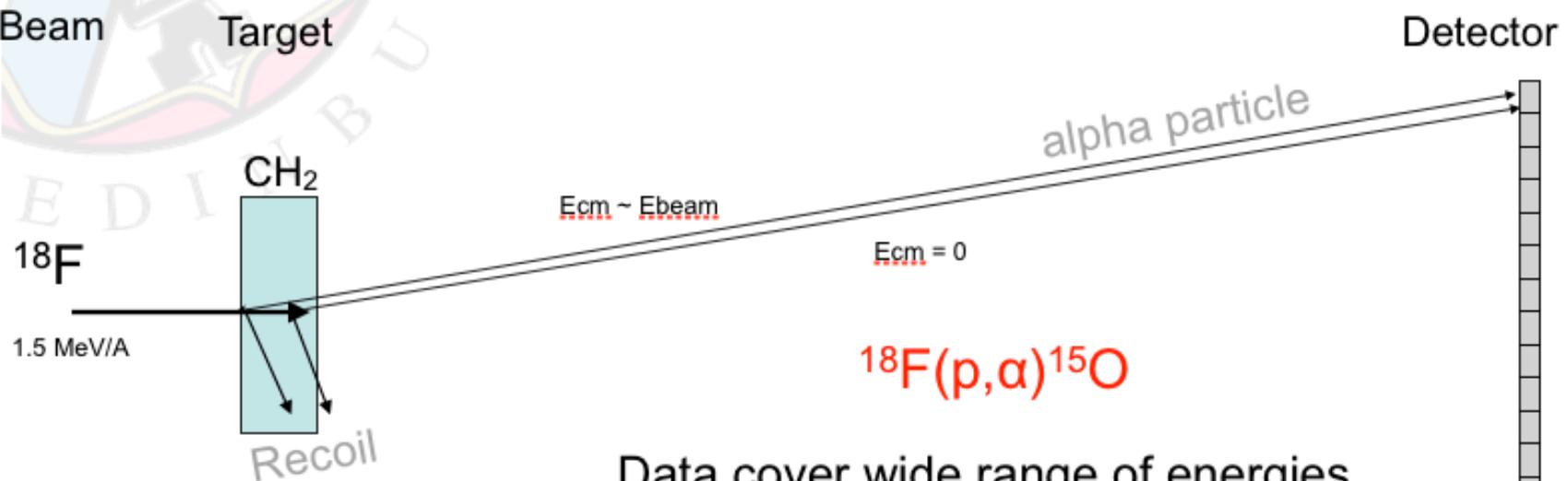


Data cover wide range of energies

Yield is combination of Coulomb scattering and resonant contribution

Detailed shape of excitation function contains required nuclear information

# Thick Target Technique

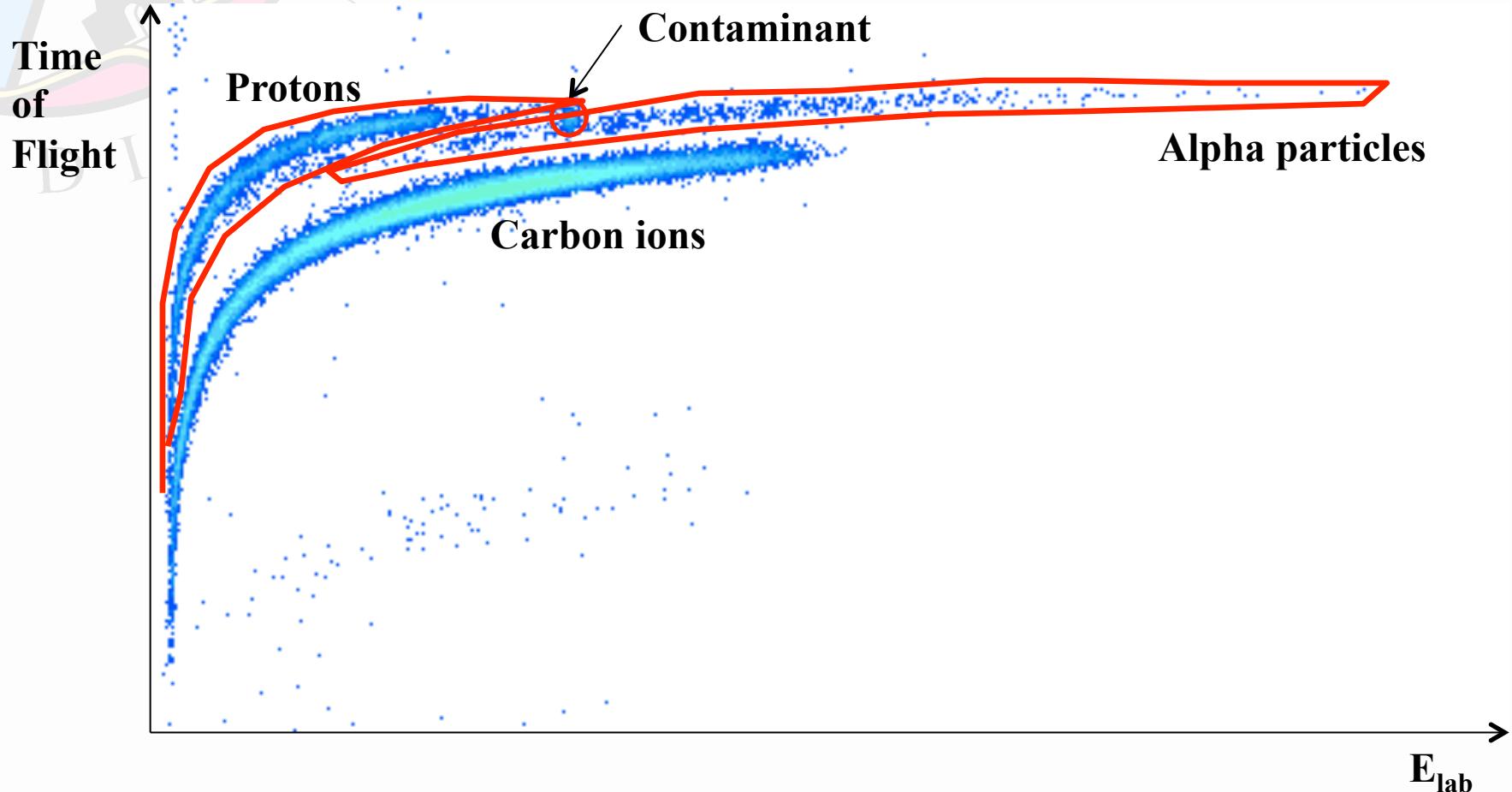


Data cover wide range of energies

Yield is from resonant contribution only

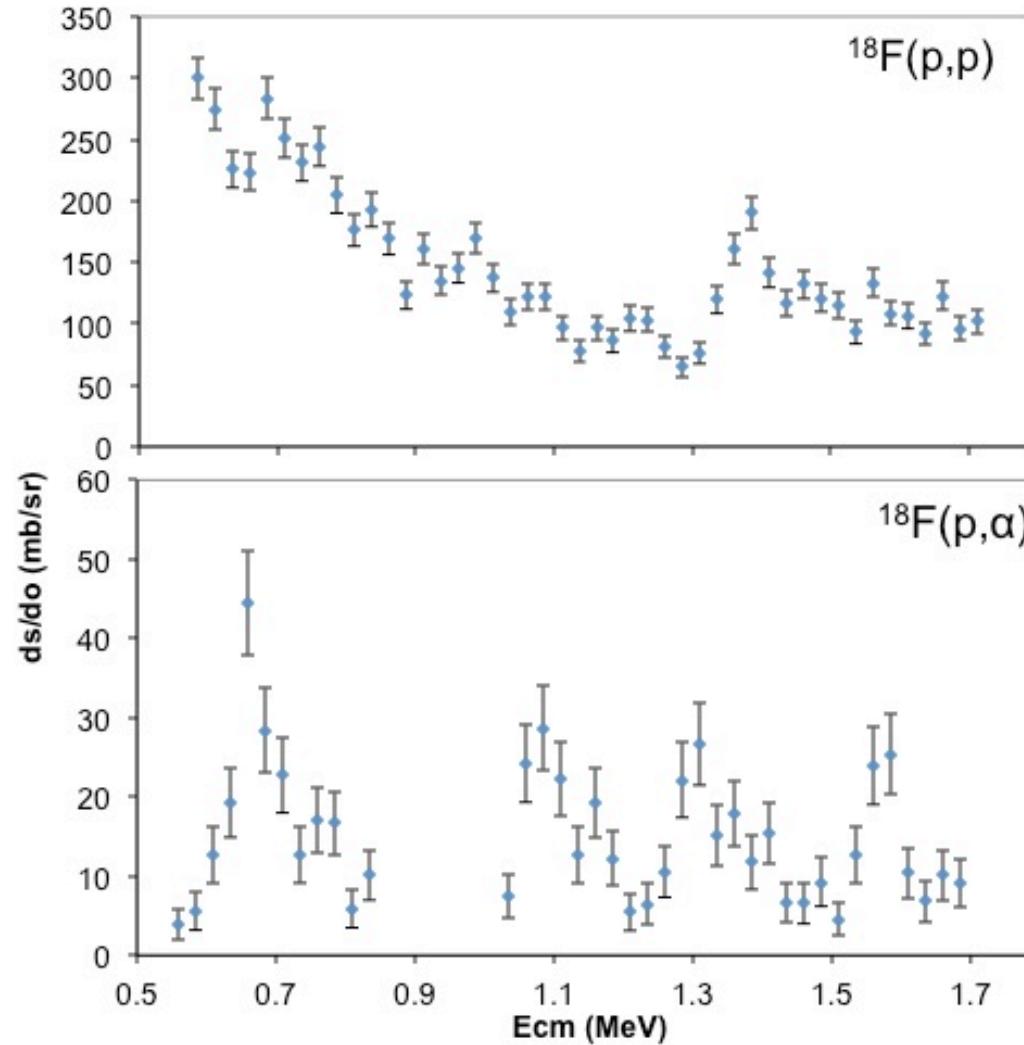
Shape is simpler

# Experimental Data





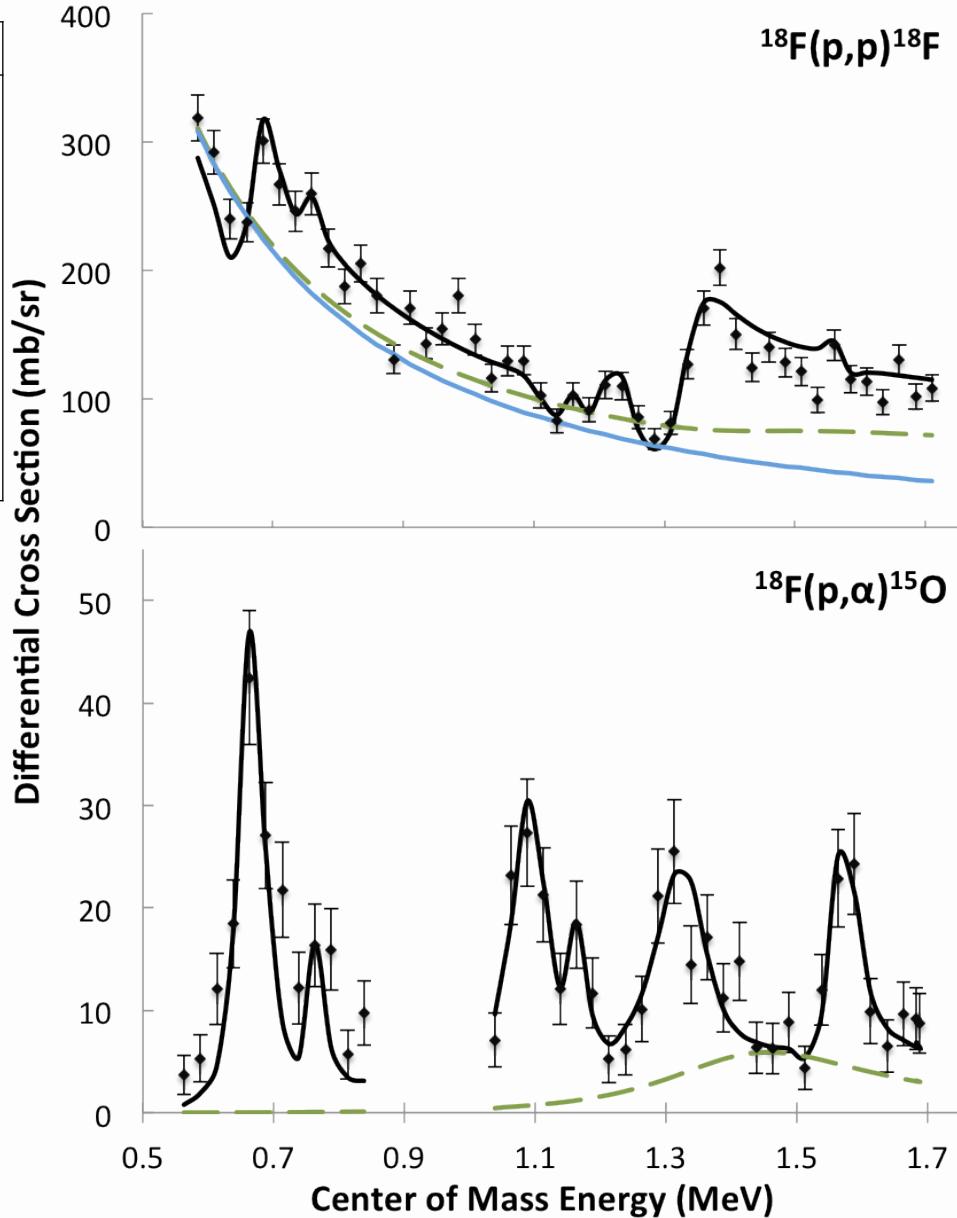
# Experimental Data



# Results

	Ecm (MeV)	$J^\pi$	$\Gamma p$ (keV)	$\Gamma \alpha$ (keV)	Int
A	0.665	$3/2^+$	15	24	
B	0.759(20)	$3/2^+$	1.6(5)	2.4(6)	
C	1.096(11)	$5/2^+$	3(1)	54(12)	
D	1.160(34)	$3/2^+$	2.3(6)	1.9(6)	
E	1.219(22)	$3/2^-$	21(3)	0.1(1)	
F	1.335(6)	$3/2^+$	65(8)	26(4)	-
G	1.455(38)	$1/2^+$	55(12)	347(92)	
H	1.571(13)	$5/2^+$	1.7(4)	12(3)	

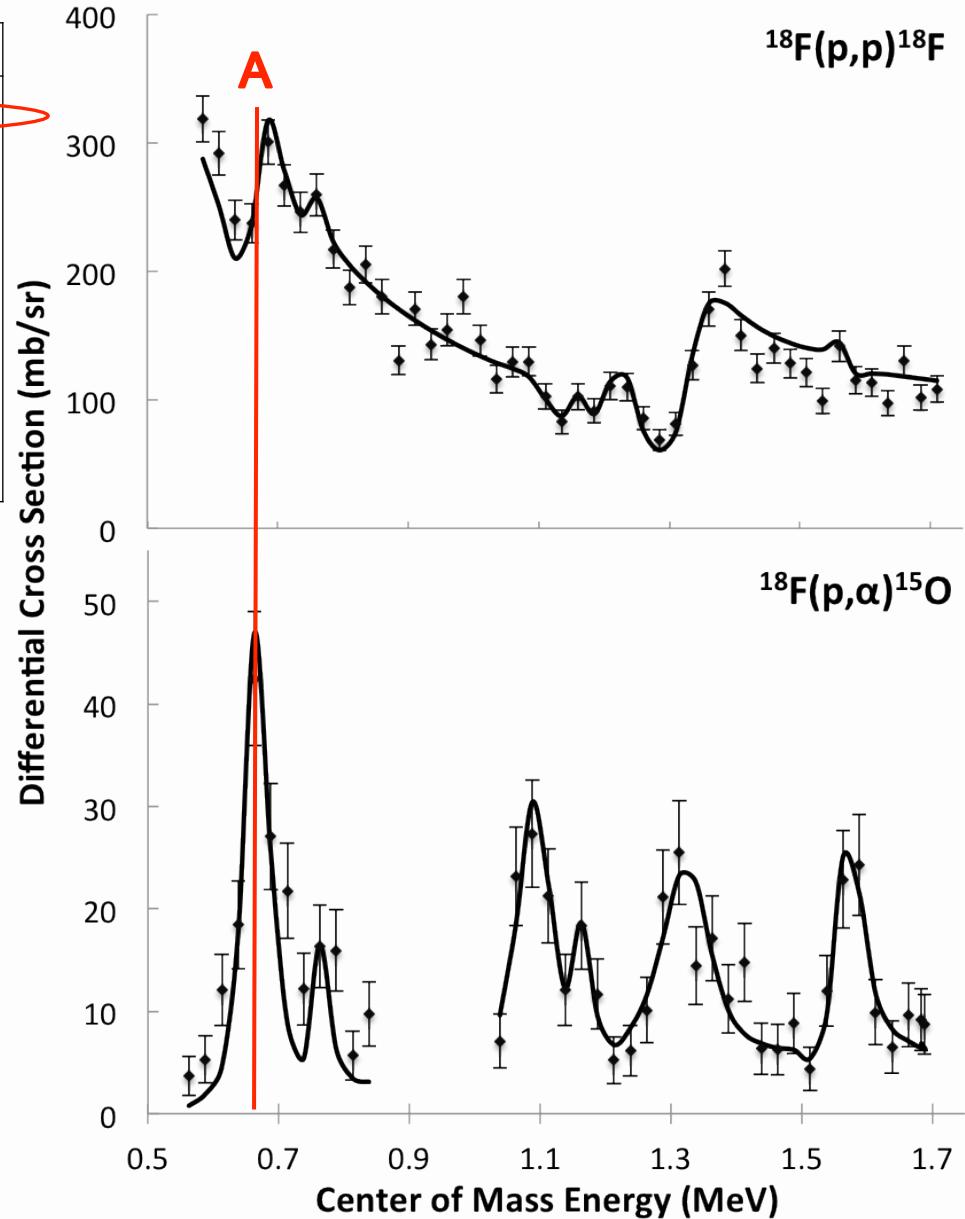
- Results of R Matrix analysis carried out using DREAM code from P. Descouvemont
- Mountford *et al.*, Phys. Rev. C **85**, 022801(R) (2012)



# Preliminary Results (A)

	Ecm (MeV)	$J^\pi$	$\Gamma p$ (keV)	$\Gamma \alpha$ (keV)	Int
A	0.665	$3/2^+$	15	24	+
B	0.759(20)	$3/2^+$	1.6(5)	2.4(6)	
C	1.096(11)	$5/2^+$	3(1)	54(12)	
D	1.160(34)	$3/2^+$	2.3(6)	1.9(6)	
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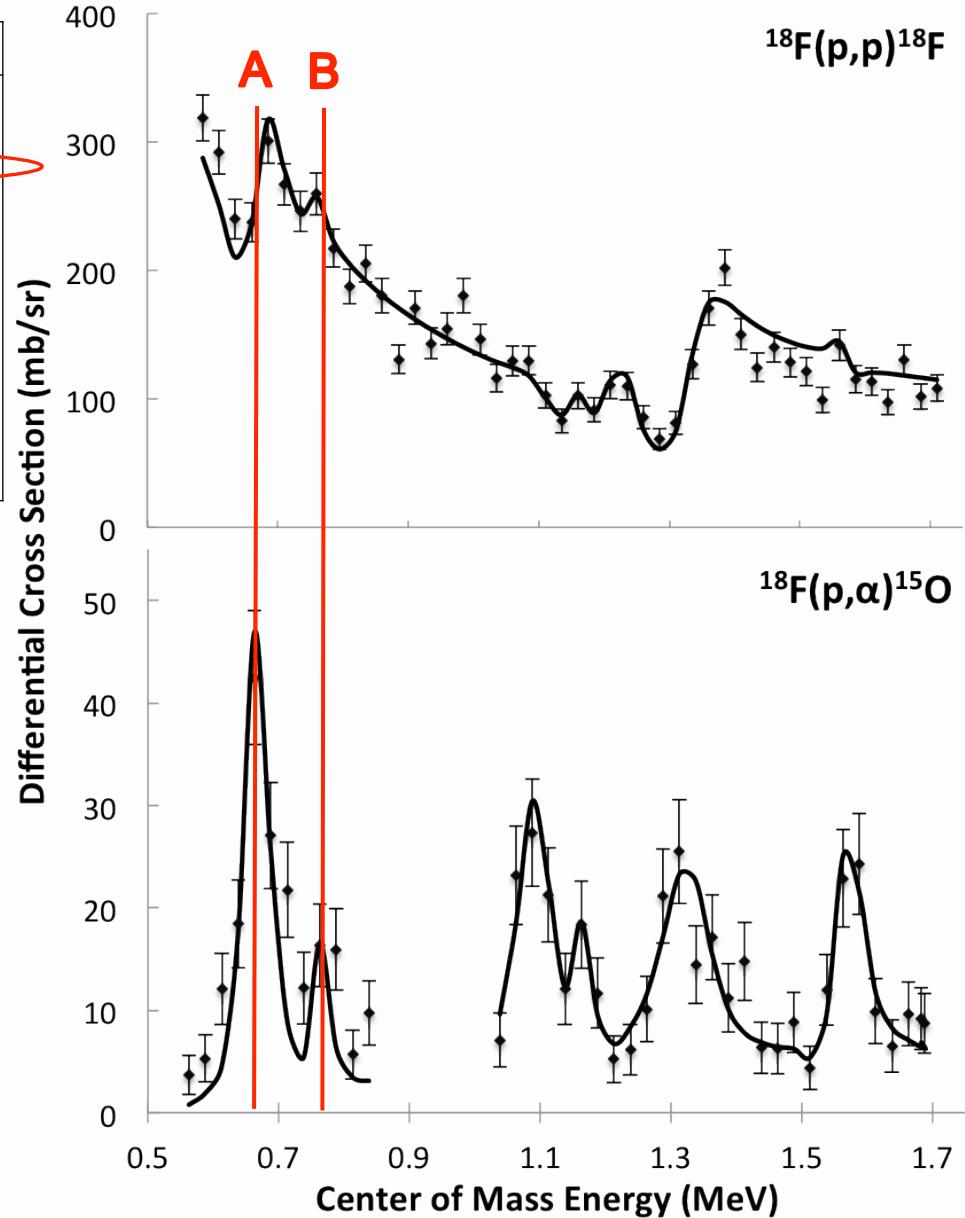
- Well known ‘665keV’ state
- Cross sections scaled to this state



# Preliminary Results (B)

	Ecm (MeV)	$J^\pi$	$\Gamma p$ (keV)	$\Gamma \alpha$ (keV)	Int
A	0.665	$3/2^+$	15	24	+
B	0.759(20)	$3/2^+$	1.6(5)	2.4(6)	
C	1.096(11)	$5/2^+$	3(1)	54(12)	
D	1.160(34)	$3/2^+$	2.3(6)	1.9(6)	
E	1.219(22)	$3/2^-$	21(3)	0.1(1)	
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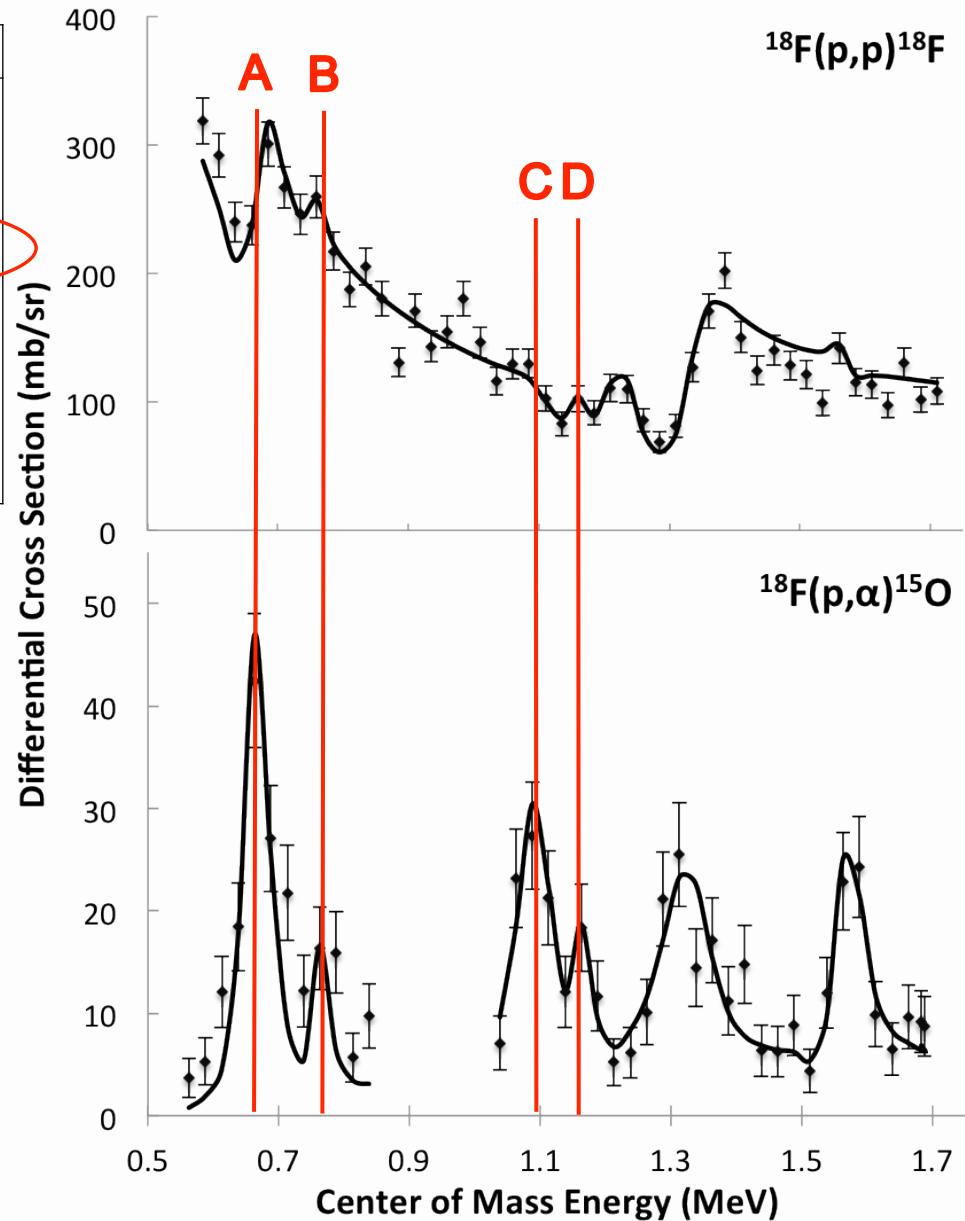
- Reported by Nesaraja/  
Dalouzy
- Significantly weaker than  
Dalouzy
- Consistent in strength  
with Nesaraja



# Preliminary Results (C & D)

	Ecm (MeV)	$J^\pi$	$\Gamma p$ (keV)	$\Gamma \alpha$ (keV)	Int
A	0.665	$3/2^+$	15	24	+
B	0.759(20)	$3/2^+$	1.6(5)	2.4(6)	
C	1.096(11)	$5/2^+$	3(1)	54(12)	
D	1.160(34)	$3/2^+$	2.3(6)	1.9(6)	
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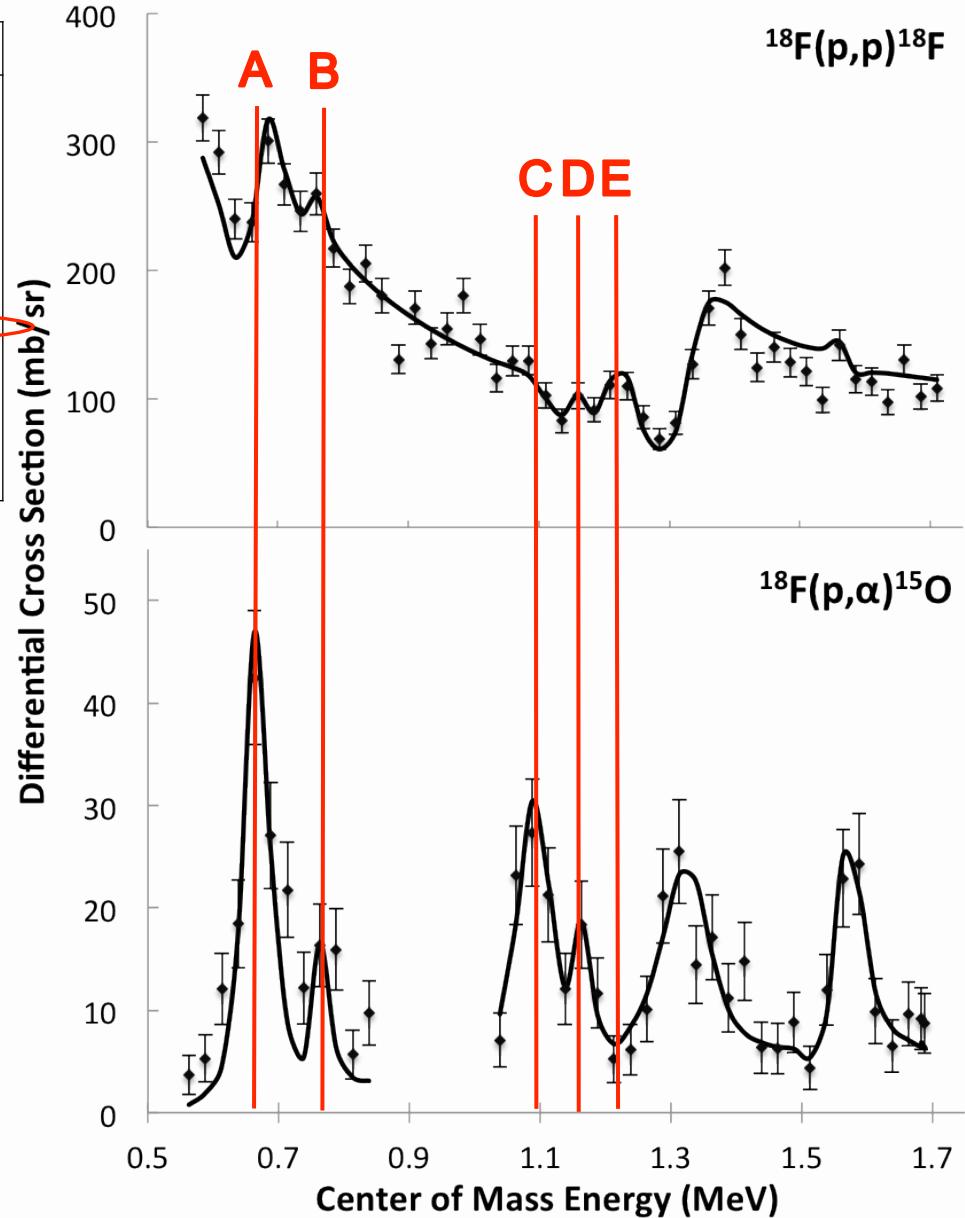
- C reported by Murphy – narrower but current result relatively consistent
- C and D reported by Dalouzy/Nesaraja at different strengths
- C in good agreement in energy



# Preliminary Results (E)

	Ecm (MeV)	$J^\pi$	$\Gamma p$ (keV)	$\Gamma \alpha$ (keV)	Int
A	0.665	$3/2^+$	15	24	
B	0.759(20)	$3/2^+$	1.6(5)	2.4(6)	
C	1.096(11)	$5/2^+$	3(1)	54(12)	
D	1.160(34)	$3/2^+$	2.3(6)	1.9(6)	
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G	1.455(38)	$1/2^+$	55(12)	347(92)	
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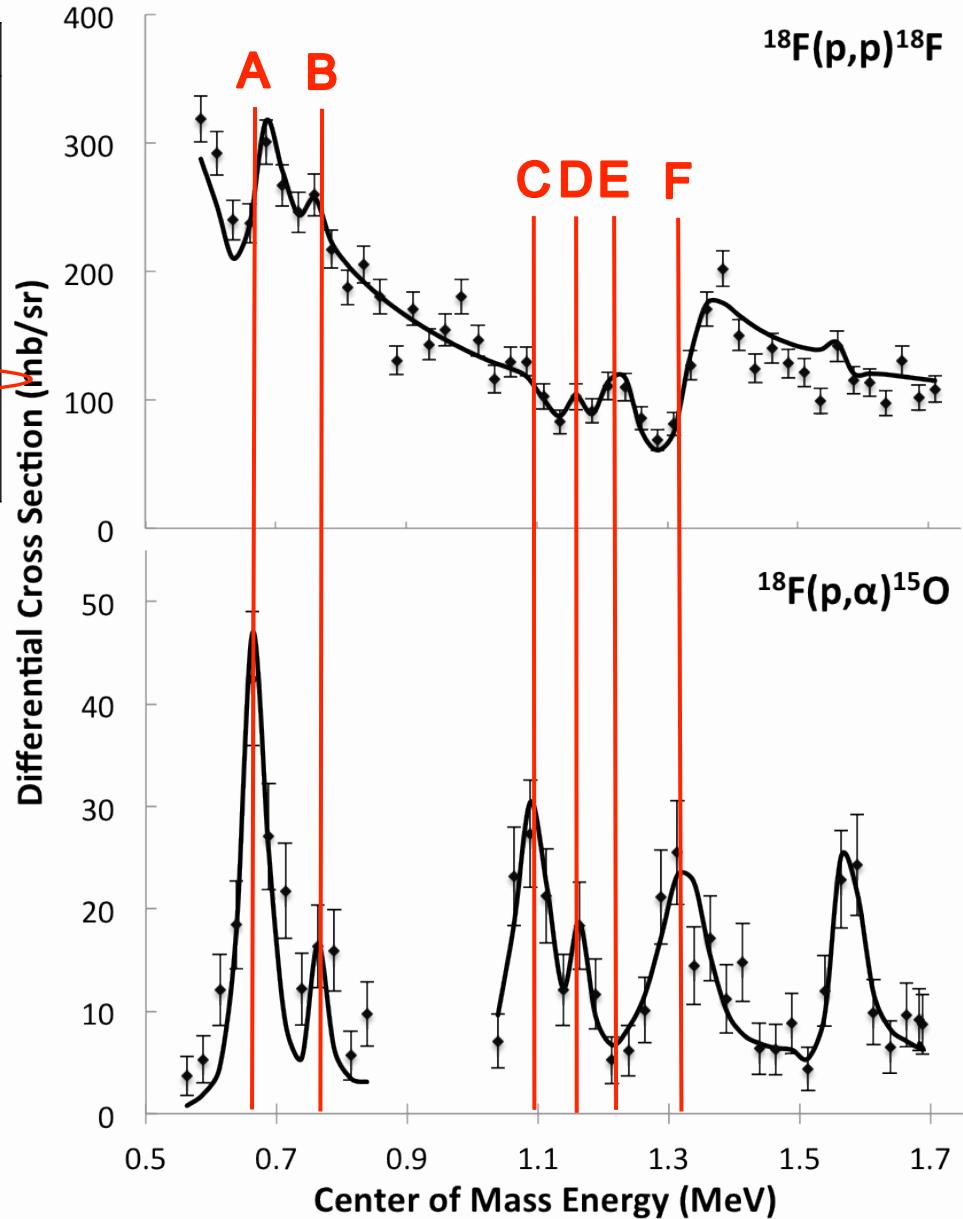
- Previously reported by Nesaraja and Murphy
- Agreement in spin with Murphy
- No agreement in strength
- Required to fit bottom of state F



# Preliminary Results (F)

	Ecm (MeV)	$J^\pi$	$\Gamma p$ (keV)	$\Gamma \alpha$ (keV)	Int
A	0.665	$3/2^+$	15	24	
B	0.759(20)	$3/2^+$	1.6(5)	2.4(6)	
C	1.096(11)	$5/2^+$	3(1)	54(12)	
D	1.160(34)	$3/2^+$	2.3(6)	1.9(6)	
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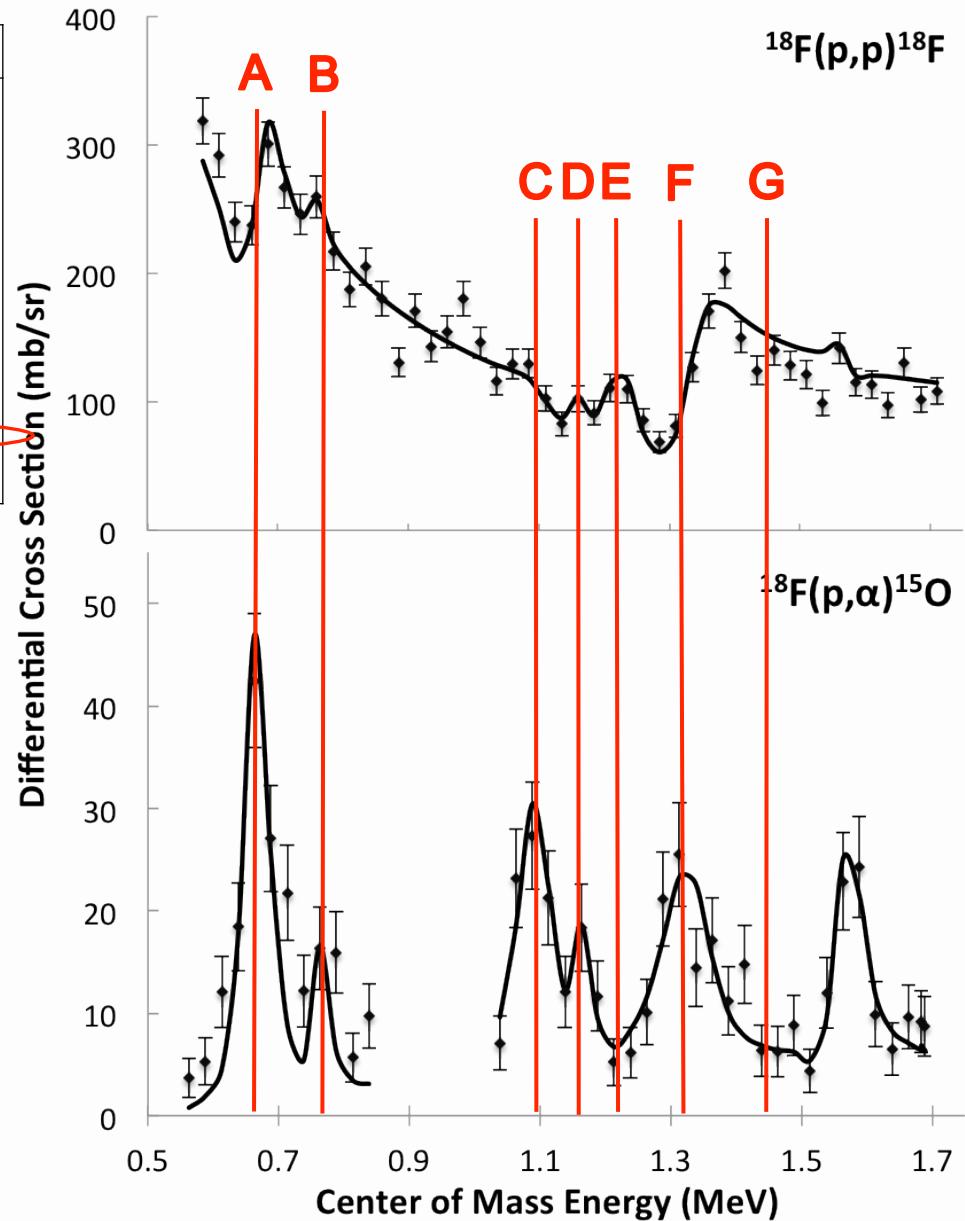
- Observed by Murphy with less strength
- New strength due to strong correlation with state G



# Preliminary Results (G)

	Ecm (MeV)	$J^\pi$	$\Gamma p$ (keV)	$\Gamma \alpha$ (keV)	Int
A	0.665	$3/2^+$	15	24	
B	0.759(20)	$3/2^+$	1.6(5)	2.4(6)	
C	1.096(11)	$5/2^+$	3(1)	54(12)	
D	1.160(34)	$3/2^+$	2.3(6)	1.9(6)	
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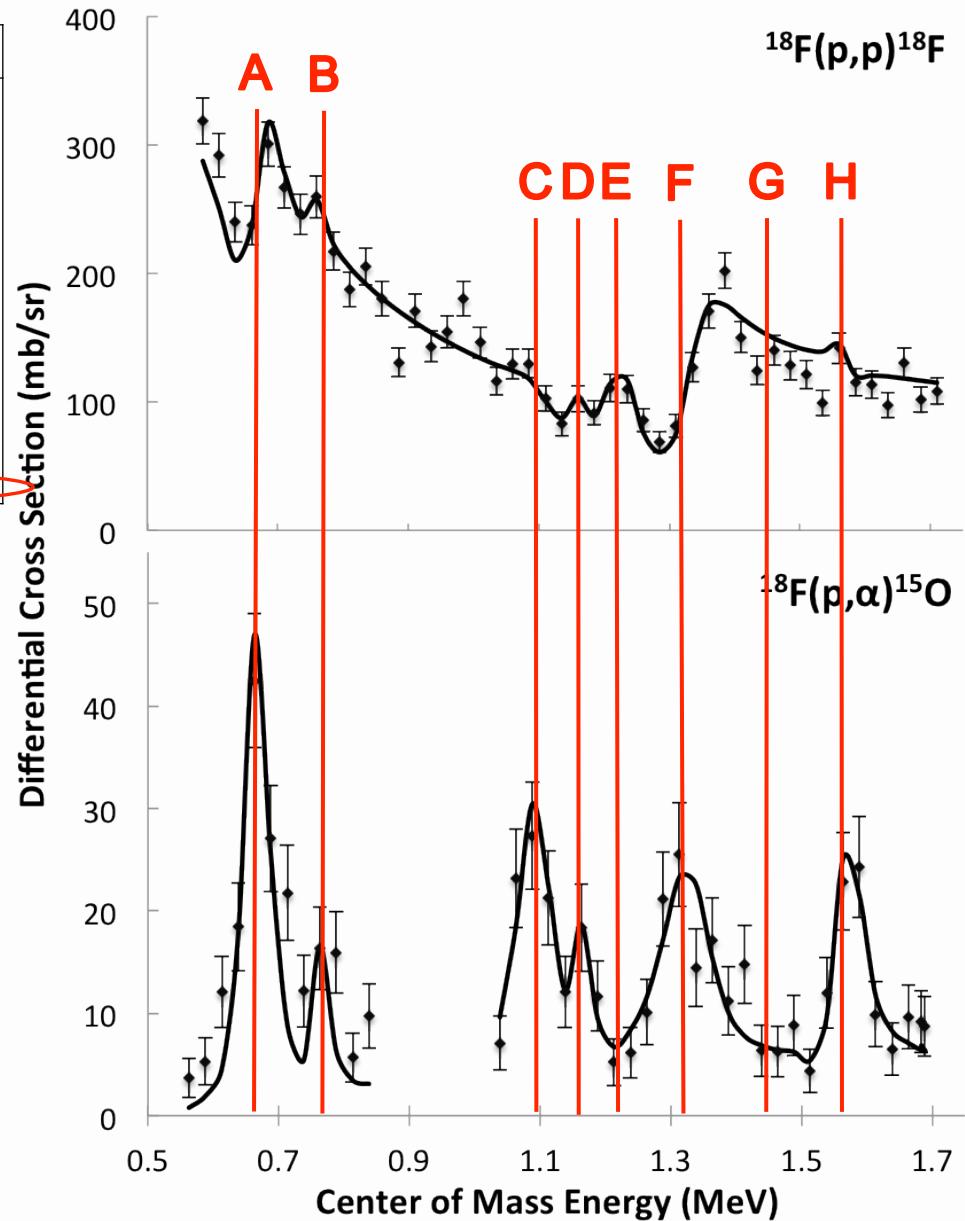
- Descouvemont state?
- Stay tuned...



# Preliminary Results (H)

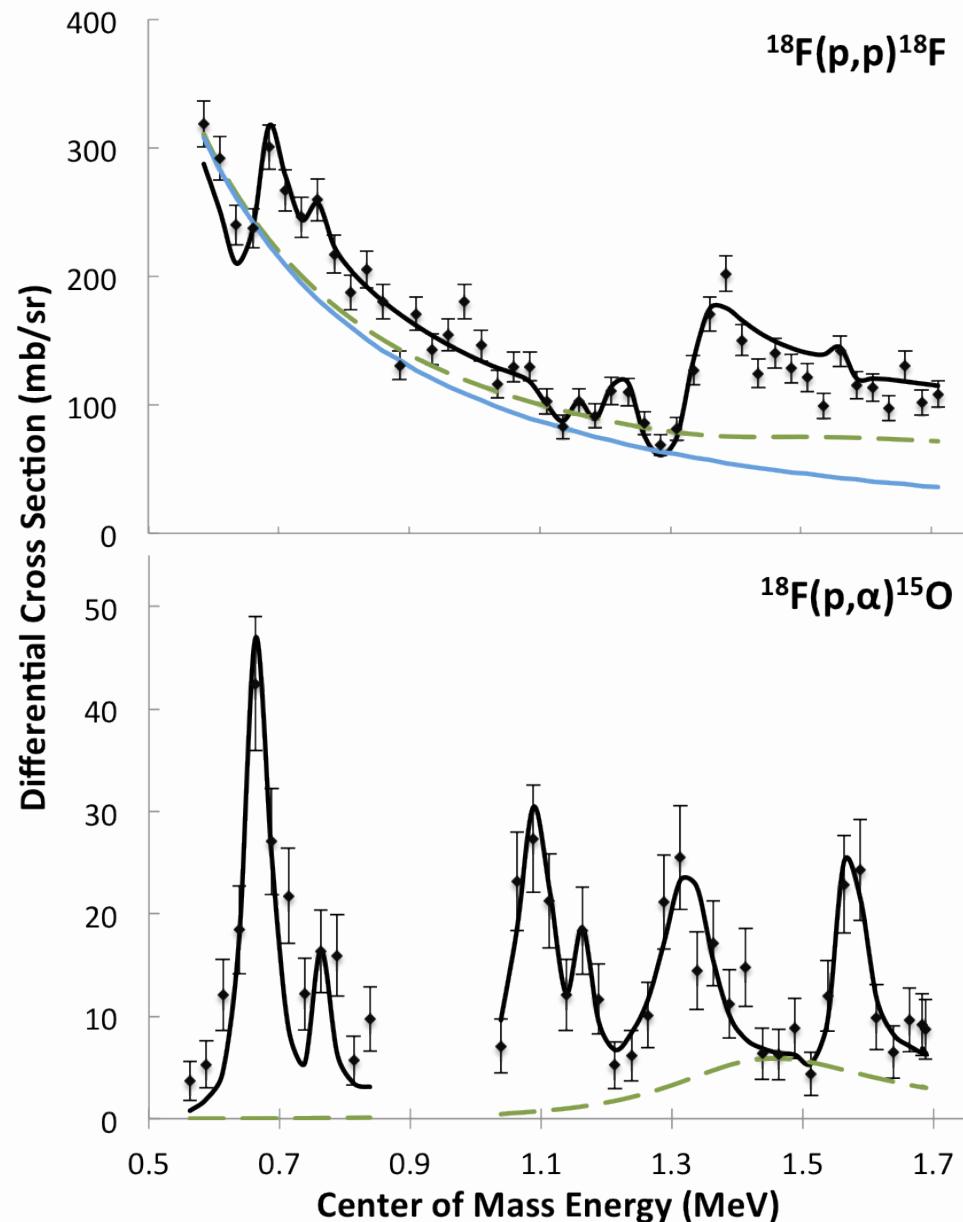
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- Observed by Dalouzy (alternative  $\pi$ ) and Murphy (alternative  $J$ )
- Dalouzy  $J$  unambiguous but  $\pi$  inferred
- Current data strongly favours  $J^\pi=5/2^+$  at consistent strengths



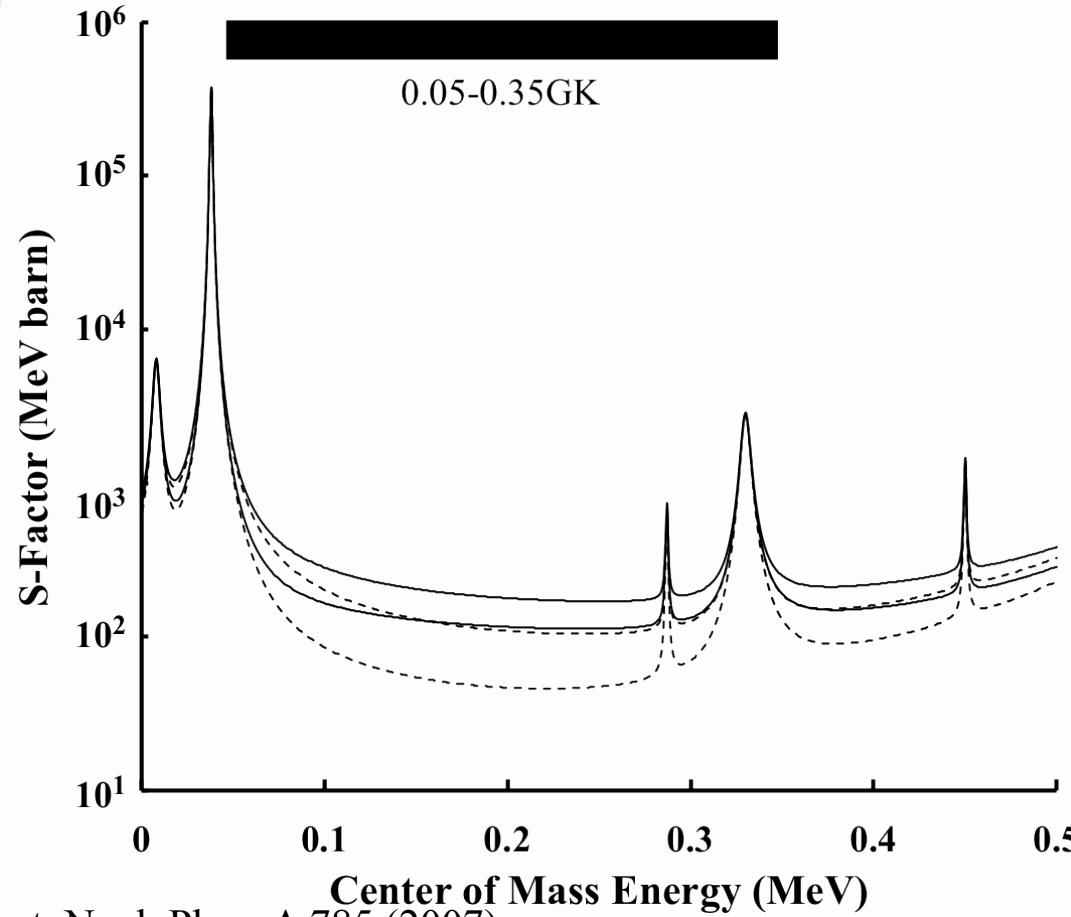
# Dufour/Descouvemont State??

- Candidate state observed at  $E_{CM}=1.455\text{MeV}$  (G)
- More than a factor of 2 *narrower* than predicted in proton channel
- More than a factor of 2 *broader* than predicted in alpha channel
- But consistent in total width ( $402(93)\text{keV}$ ) with Dalouzy ( $292(107)\text{keV}$ )
- Broad state is required to fit to data



# Impact: Enhanced S-Factor

- Low energy and sub-threshold parameters as in [3]
- Enhanced by prediction of [2]



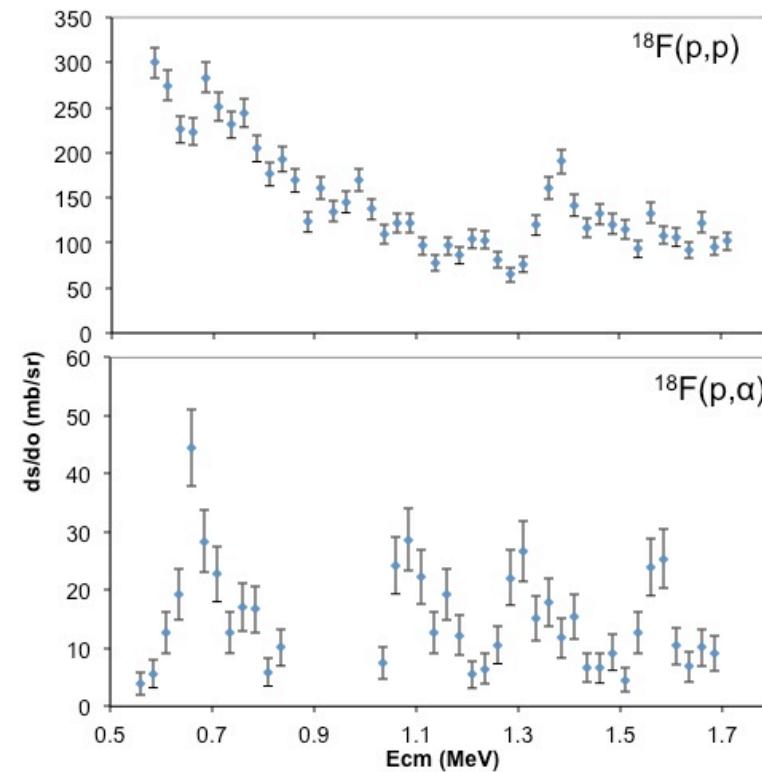
[2] Dufour/Descouvemont, Nucl. Phys. A 785 (2007)

[3] Iliadis *et al.*, Nucl. Phys. A 841, 251 (2010)

# A bit more about the fitting...

## Problems

- Many parameter fit
- 10 x statistics in  $(p,p)$  channel compared to  $(p,\alpha)$
- Changing energy resolution





## Energy Resolution

## Complexity

## Statistics

# Solutions...

- Monte Carlo to estimate energy resolution
- $dE(E)$  in (some) R-Matrix code
- Iterative analysis
  - Literature values
  - Trial and error (by eye adjustments)
  - Fitting single resonances
  - Fitting multiple resonances
  - Fitting multiple channels
  - Fitting entire data set
- Play God
  - Initially adjust error bars to change weightings
  - Make active use of error matrix (covariance)

Include additional contributions to  
final error bars!

# Co-variance

TABLE II. Covariance matrix for all parameters allowed to vary in the fitting process.

# Alternative R-Matrix Implementation: AZURE

- Generally available code from JINA
- No limit to number of channels

$$\frac{d\sigma}{d\Omega} \propto |U_J U_{J'}^*|$$

PHYSICAL REVIEW C 81, 045805 (2010)

## AZURE: An *R*-matrix code for nuclear astrophysics

R. E. Azuma,<sup>1,2</sup> E. Uberseder,<sup>2,\*</sup> E. C. Simpson,<sup>2,3</sup> C. R. Brune,<sup>4</sup> H. Costantini,<sup>2,5</sup> R. J. de Boer,<sup>2</sup> J. Görres,<sup>2</sup> M. Heil,<sup>6</sup> P. J. LeBlanc,<sup>2</sup> C. Ugalde,<sup>2,†</sup> and M. Wiescher<sup>2</sup>

<sup>1</sup>*Department of Physics, University of Toronto, Toronto, Ontario M5S 1A7, Canada*

<sup>2</sup>*University of Notre Dame, Department of Physics, Notre Dame, Indiana 46556, USA*

<sup>3</sup>*Department of Physics, Faculty of Engineering and Physical Sciences, University of Surrey, Guildford, Surrey GU2 7XH, United Kingdom*

<sup>4</sup>*Department of Physics and Astronomy, Ohio University, Ohio 45701, USA*

<sup>5</sup>*Istituto Nazionale Fisica Nucleare (INFN), Genova, Italy*

<sup>6</sup>*GSI Darmstadt, Planckstr. 1, 64291 Darmstadt, Germany*

(Received 11 January 2010; published 26 April 2010)

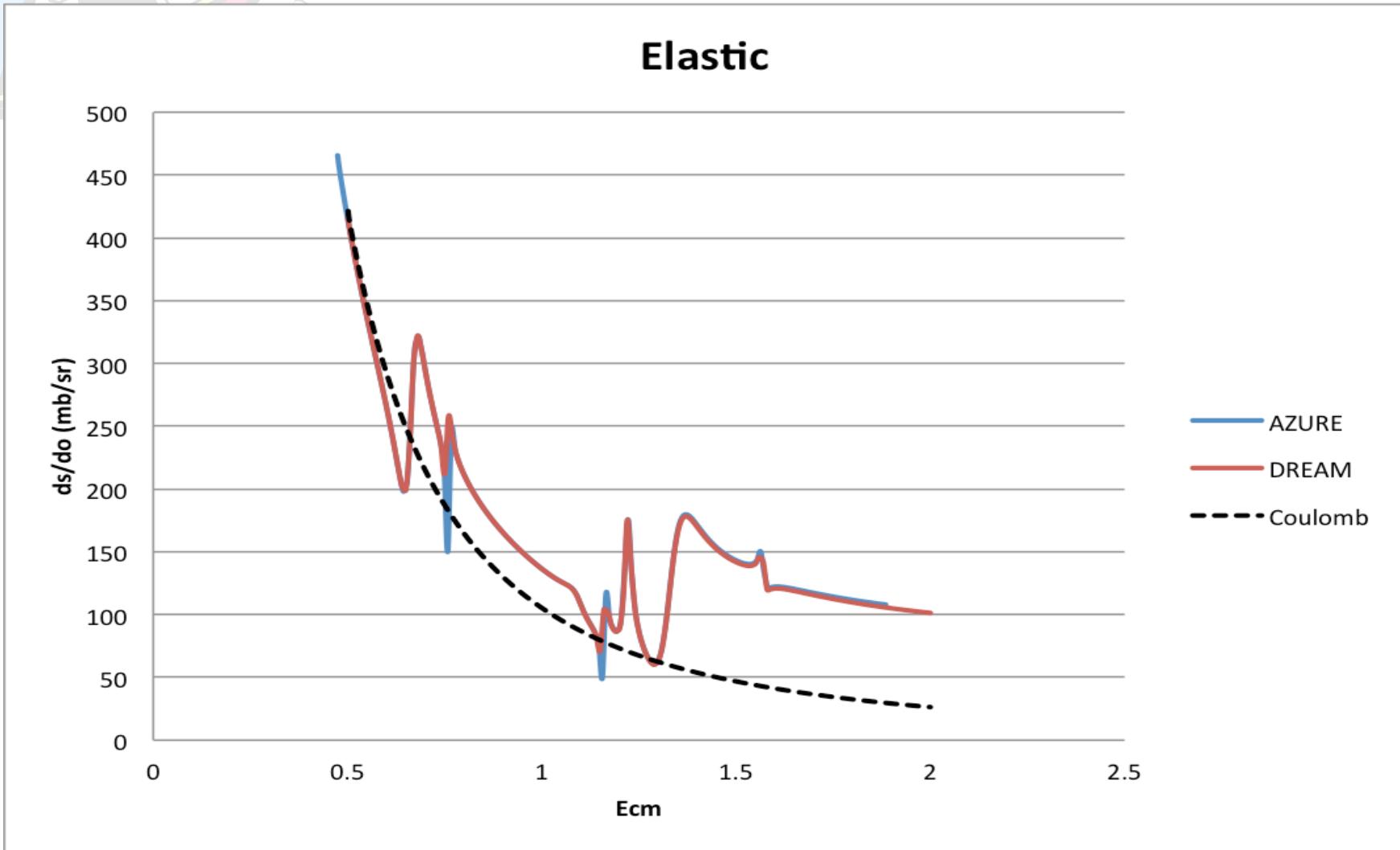
The paper describes a multilevel, multichannel *R*-matrix code, AZURE, for applications in nuclear astrophysics. The code allows simultaneous analysis and extrapolation of low-energy particle scattering, capture, and reaction cross sections of relevance to stellar hydrogen, helium, and carbon burning. The paper presents a summary of *R*-matrix theory, code description, and a number of applications to demonstrate the applicability and versatility of AZURE.

DOI: [10.1103/PhysRevC.81.045805](https://doi.org/10.1103/PhysRevC.81.045805)

PACS number(s): 26.20.Cd, 25.40.Lw, 24.10.-i

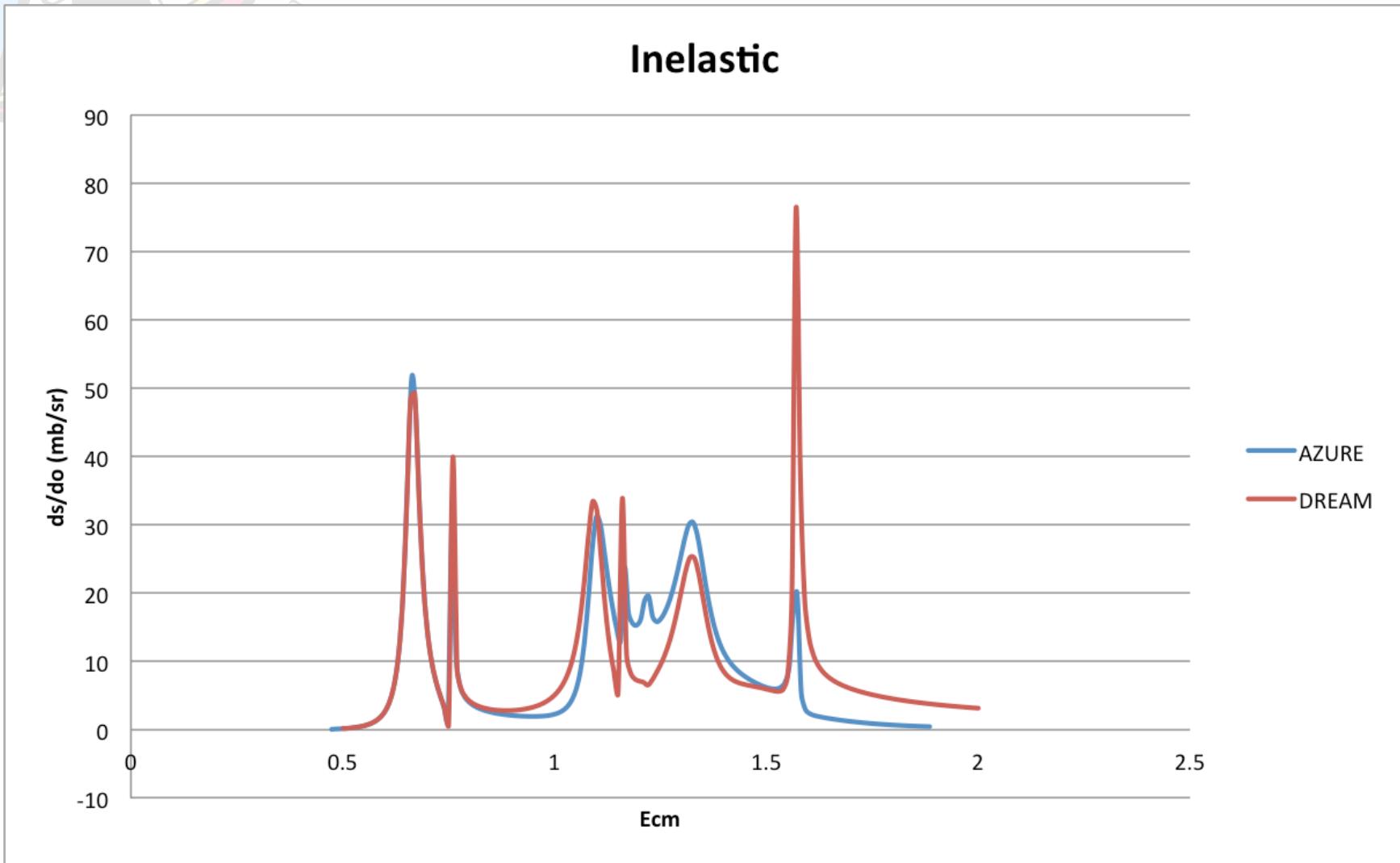
# R-Matrix Comparison

- Same parameters in each code:



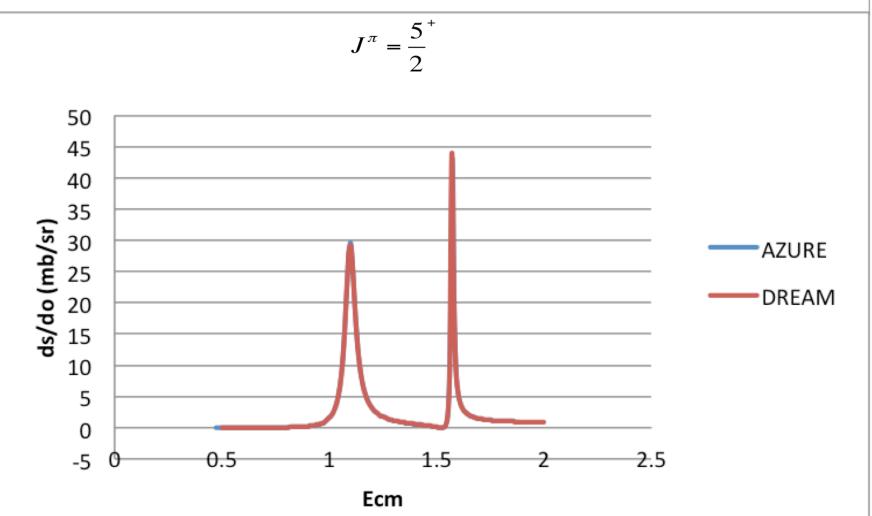
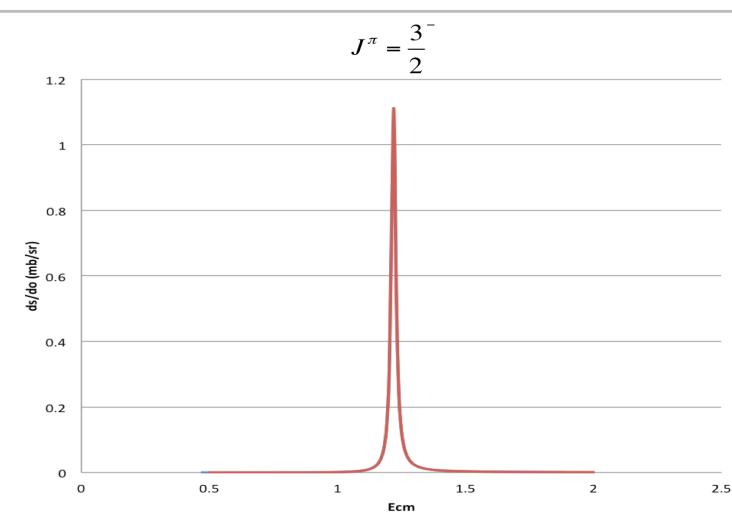
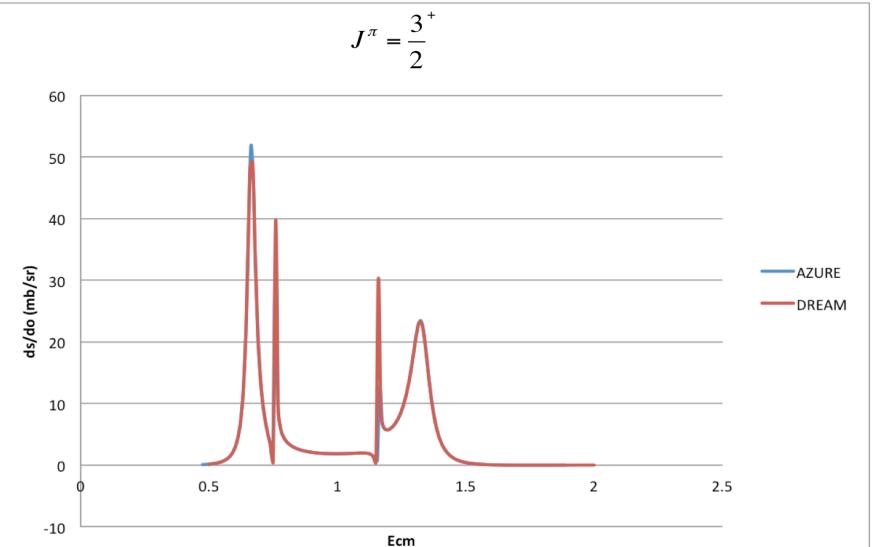
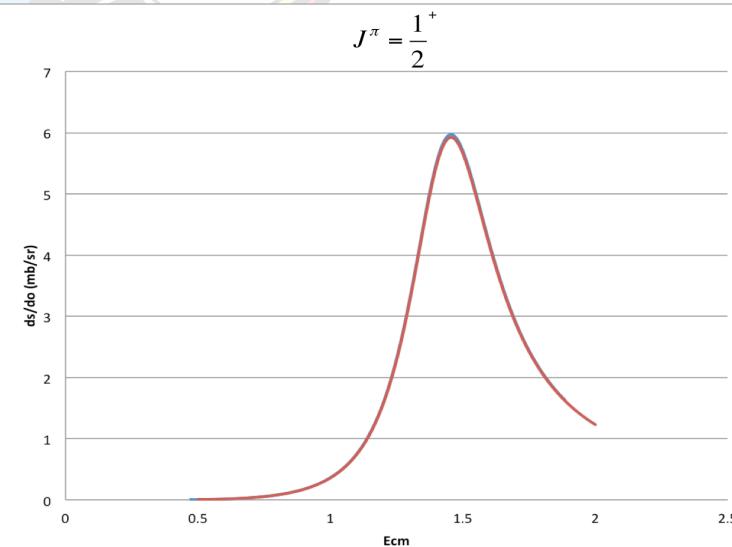
# R-Matrix Comparison

- Same parameters in each code:



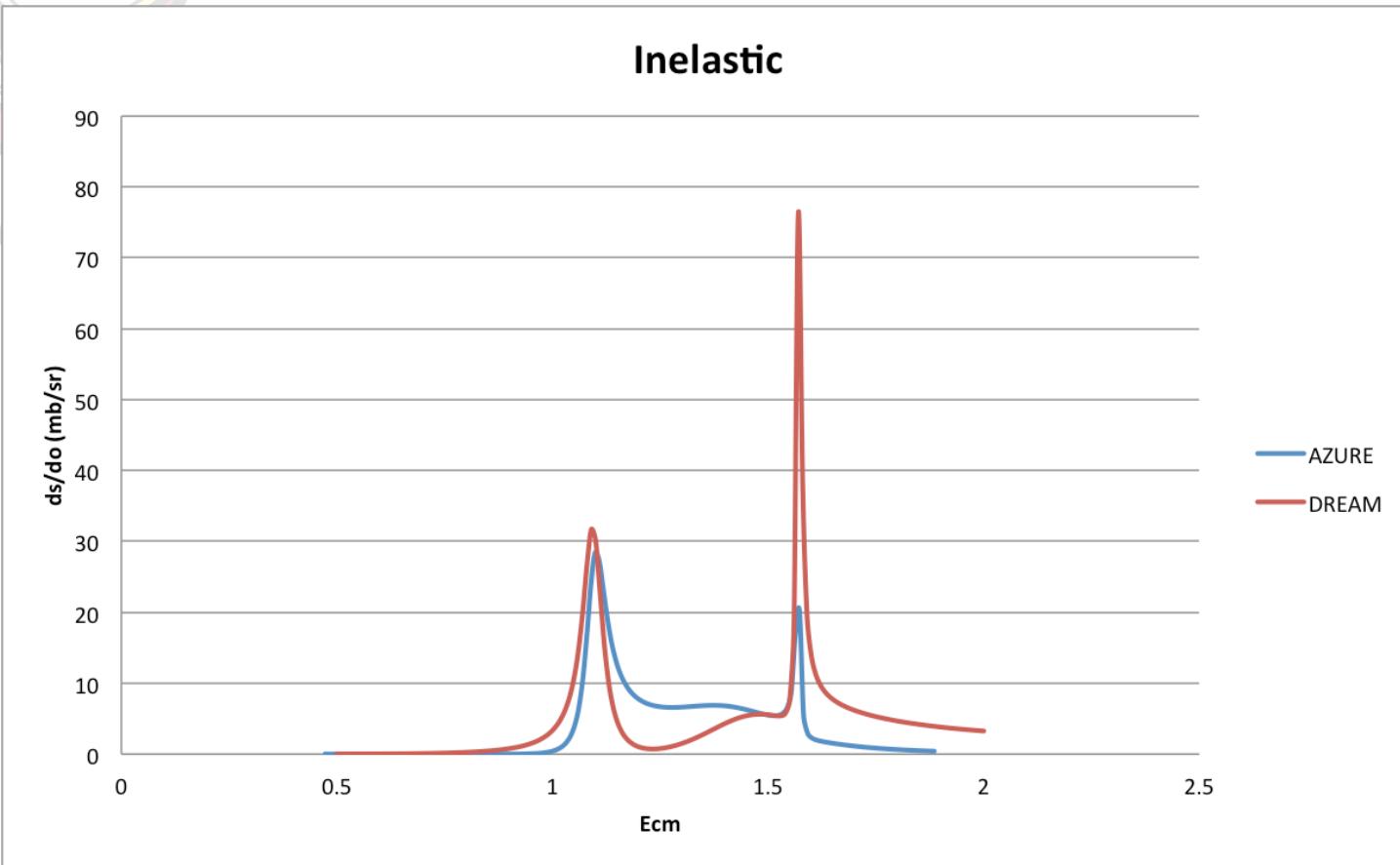
# R-Matrix Comparison

- Individual  $J^\pi$  groups:



# R-Matrix Comparison

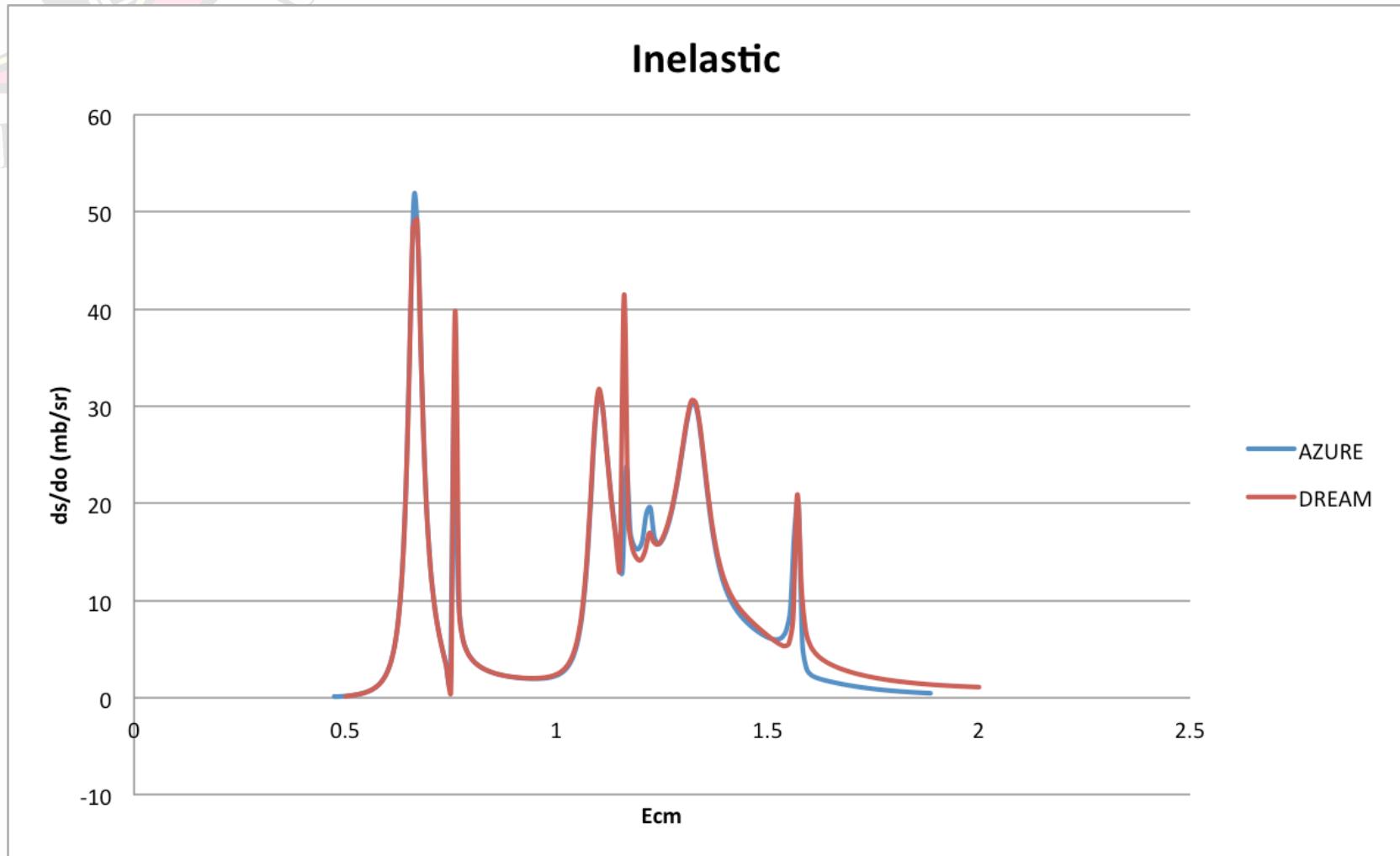
- Combining  $1/2^+$  and  $5/2^+$ :



- Opposite interference between groups
- Discrepancy is not in angular coupling coefficients

# R-Matrix Comparison

- Discrepancy arises in use of arctan to calculate  $\delta$ 's, alternative interference:



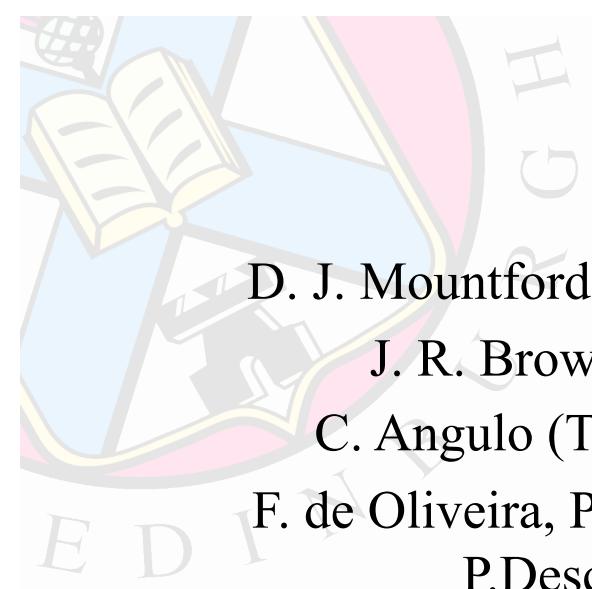
# R-Matrix Comparison

Mountford et al [6]				AZURE Results		
$E_{CM}$ (MeV)	$J^\pi$	$\Gamma_p$ (keV)	$\Gamma_\alpha$ (keV)	$E_{CM}$ (MeV)	$\Gamma_p$ (keV)	$\Gamma_\alpha$ (keV)
0.665	$\frac{3}{2}^+$	15.2	23.8	0.665	15.2	23.8
0.759(20)	$\frac{3}{2}^+$	1.6(5)	2.4(6)	0.755(77)	1.7(17)	3.3(33)
1.096(11)	$\frac{5}{2}^+$	3(1)	54(12)	1.097(29)	3.3(20)	71(45)
1.160(34)	$\frac{3}{2}^+$	2.3(6)	1.9(6)	1.149(14)	2.5(25)	2.1(21)
1.219(22)	$\frac{3}{2}^-$	21(3)	0.1(1)	1.211(17)	27(17)	0.2(2)
1.335(6)	$\frac{3}{2}^+$	65(8)	26(4)	1.339(23)	65(13)	31(27)
1.455(38)	$\frac{1}{2}^+$	55(12)	347(92)	1.498(176)	44(23)	313(147)
1.571(13)	$\frac{5}{2}^+$	1.7(4)	12(3)	1.569(30)	1.7(12)	8(6)

# Conclusions

- R-matrix formalism extremely useful for extracting resonance parameters for astrophysical reactions
- New data obtained in study of astrophysically important  $^{18}\text{F}(\text{p},\alpha)^{15}\text{O}$  reaction
- Analysis, aided by R-Matrix calculations, finds candidate for newly proposed s-wave state

**Enhanced rate of  $^{18}\text{F}$  destruction → less  $^{18}\text{F}$   
→ detectability distance reduced**



# Collaborators

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# Thank You