# Reaction tools for the study of exotic nuclei: theory and applications - I

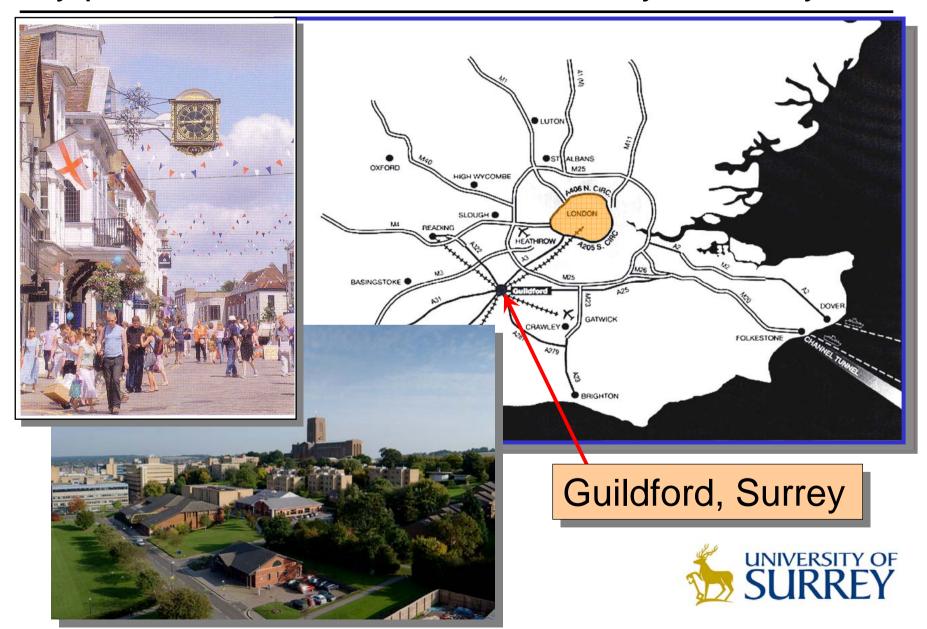
Hadrons and Nuclei under Extreme Conditions (HANEC2010),

TIT, O-Okayama

16-17 September 2010

Jeff Tostevin, TIT and Department of Physics University of Surrey, UK

### My part of the Cosmos - the University of Surrey



#### Exotic nuclei – fundamental questions

What are the limits of nuclear existence ...?

How do the limits of weak binding and extreme proton-to-neutron asymmetries affect nuclear properties ...?

How do the properties of nuclei evolve with changes in proton and neutron number ...?

"Atomic Nuclei: Structure and Stability", US 2002 Long-Range Plan for Nuclear Science

Some other slides/notes/resources can be found at:

http://www.nucleartheory.net/DTP\_material/

#### There several good direct reactions texts:

<u>Direct nuclear reaction theories</u> (Wiley, Interscience monographs and texts in physics and astronomy, v. 25) <u>Norman Austern</u>

<u>Direct Nuclear Reactions</u> (Oxford University Press, International Series of Monographs on Physics, 856 pages ) <u>G R Satchler</u>

Introduction to the Quantum Theory of Scattering (Academic, Pure and Applied Physics, Vol 26, 398 pages) LS Rodberg, RM Thaler

<u>Direct Nuclear Reactions</u> (World Scientific Publishing, 396 pages)

<u>Norman K. Glendenning</u>

<u>Introduction to Nuclear Reactions</u> (Taylor & Francis, Graduate Student Series in Physics, 515 pages ) <u>C A Bertulani, P Danielewicz</u>

<u>Theoretical Nuclear Physics: Nuclear Reactions</u> (Wiley Classics Library, 1938 pages ) <u>Herman Feshbach</u>

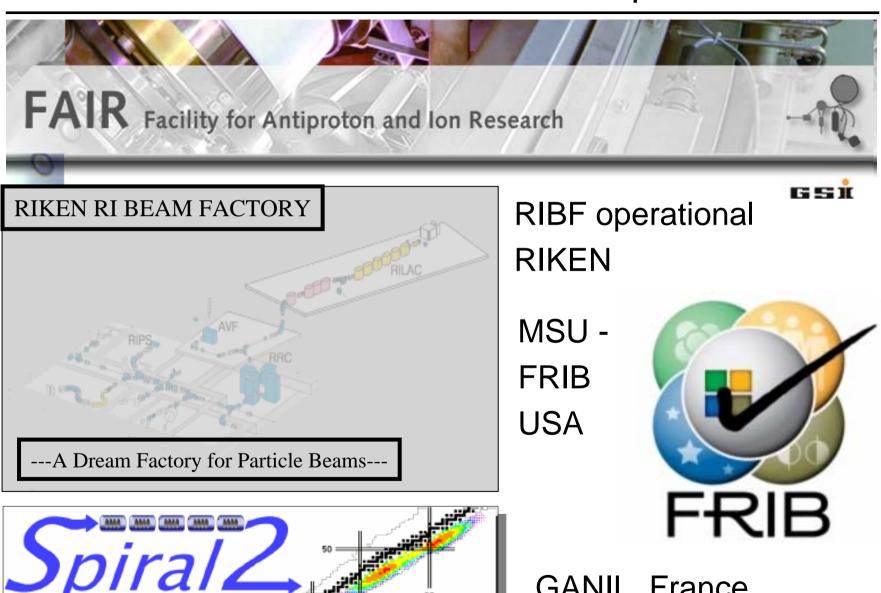
Introduction to Nuclear Reactions (Oxford University Press, 332 pages)

G R Satchler

Nuclear Reactions for Astrophysics (Cambridge University Press, 2010)

Ian Thompson and Filomena Nunes

#### Radioactive ion-beam facilities – and plans



GANIL, France

#### Lectures plan

### Lecture 1: Introduction: history and key ideas

- nuclear shell structure, old and new Reactions at high-energy (removal)
- approximations, simplifications
- connection/interface with nuclear structures
- observables and what they tell us

# Lecture 2: Illustrative and topical applications

- determining exotic structures
- testing of structure models and of shell model effective interactions
- nucleon and pair 'correlations'

#### Introduction: Motivations and context

#### Exotic nuclei – stability → structures → reactions

How? history: have we got to present point?

What? are exotic nuclei – how are they different?

Why? are they of interest, and to whom?

How? (and where) are they produced

Which? can we/will we produce (in 2010-2015)?

What? are we now able to do with them?

What? are are they telling us?

Where? is this research area going/challenges

#### 100 years (almost) of nuclear physics

[ 669 ]

LXXIX. The Scattering of α and β Particles by Matter and the Structure of the Atom. By Professor E. RUTHERFORD, F.R.S., University of Manchester\*.

§ 1. IT is well known that the  $\alpha$  and  $\beta$  particles suffer deflexions from their rectilinear paths by encounters with atoms of matter. This scattering is far more marked for the  $\beta$  than for the  $\alpha$  particle on account of the much smaller momentum and energy of the former particle. There seems to be no doubt that such swiftly moving particles pass through the atoms in their path, and that the deflexions observed are due to the strong electric field traversed within the atomic system. It has generally been

Philosophical Magazine, volume 21 (1911), pages 669-688

#### Nuclear sizes: charge and mass distributions

REVIEWS OF MODERN PHYSICS

VOLUME 30, NUMBER 2

APRIL, 1958

# International Congress on Nuclear Sizes and Density Distributions

Held at Stanford University, December 17-19, 1957

REVIEWS OF MODERN PHYSICS

VOLUME 30, NUMBER 2

APRIL, 1958

# Nuclear Radii as Determined by Scattering of Neutrons

S. Fernbach

University of California Radiation Laboratory, Livermore, California

REVIEWS OF MODERN PHYSICS

VOLUME 30, NUMBER 2

APRIL, 1958

#### Nuclear Density Distributions from Proton Scattering

A. E. GLASSGOLD

Department of Physics, University of California, Berkeley, California

REVIEWS OF MODERN PHYSICS

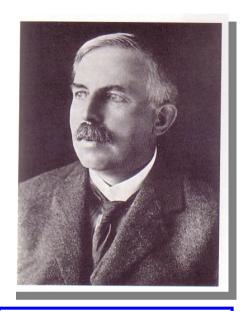
VOLUME 30, NUMBER 2

APRI

#### Electron Scattering and Nuclear Charge Distributions

D. G. RAVENHALL

Department of Physics, University of Illinois



$$\mathbf{R} = \mathbf{r}_0 \; \mathbf{A}^{1/3}$$

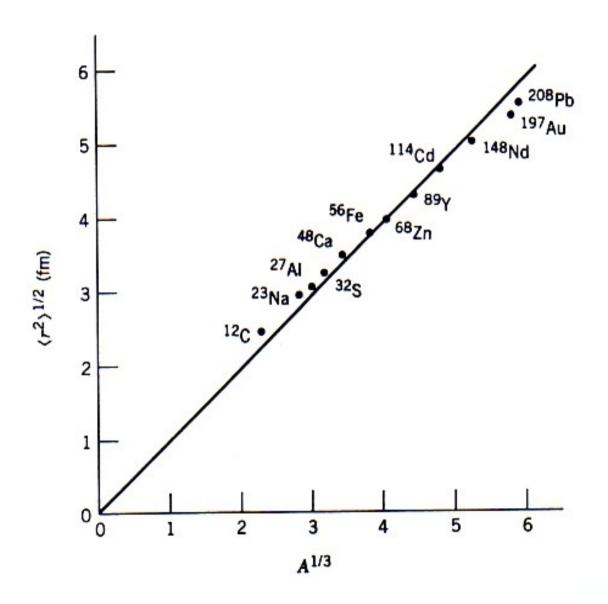
while for convenience the results on the other nucle were analyzed in terms of the Fermi shape

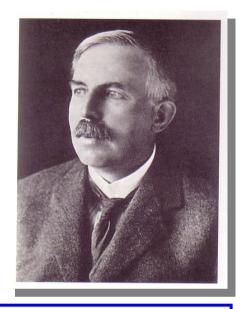
$$\rho(r) = \rho_0 [1 + e^{(r-c)/z_1}]^{-1},$$

the only significant quantities determined were the radius and the surface thickness t. The radial parameter c, the distance to the point where  $\rho$  has dropped to half of its central value on this model, varies closely as  $A^{\frac{1}{3}}$  for all of the nuclei, and t is effectively constant:

$$c = (1.07 \pm 0.02) A^{\frac{1}{3}} \times 10^{-13} \text{ cm}$$
 charge   
  $t = (2.4 \pm 0.3) \times 10^{-13} \text{ cm}$  distribution.

#### Nuclear sizes: charge and mass distributions





$$\mathbf{R} = \mathbf{r}_0 \; \mathbf{A}^{1/3}$$

r convenience the results on the other nuclearlyzed in terms of the Fermi shape

$$\rho(r) = \rho_0 [1 + e^{(r-c)/z_1}]^{-1},$$

r significant quantities determined were the ad the surface thickness t. The radial parameter istance to the point where  $\rho$  has dropped to the central value on this model, varies closely all of the nuclei, and t is effectively constant:

= 
$$(1.07\pm0.02)A^{\frac{1}{2}}\times10^{-13}$$
 cm charge  
 $t=(2.4\pm0.3)\times10^{-13}$  cm distribution.

### Shell (single particle) structures in nuclei

#### On Closed Shells in Nuclei, II

MARIA GOEPPERT MAYER

Argonne National Laboratory and Department of Physics. University of Chicago, Chicago, Illinois

February 4, 1949

Phys. Rev. 75, 1969 (1949)

Cited 240 times

LETTERS TO THE EDITOR

with spin 3/2 in stead of the expected  $d_{5/2}$ , and  $d_{25} Mn^{25}$  with 5/2 instead of the expected  $f_{7/2}$ , are the only violations.

Table II lists the known spins and orbital assignments from magnetic moments' when these are known and unambiguous, for the even-odd nuclei up to 83. Beyond 83 the data is

limited and no exceptions to the assignment appear.

Up to Z or N=20, the assignment is the same as that of Feenberg and Nordheim. At the beginning of the next shell, for levels occur at 21 and 23, as they should. At 28 the fre levels should be filled, and no spins of 7/2 are encountered any more in this shell. This subshell may contribute to the stability of Ca<sup>48</sup>. If the  $g_{2/2}$  level did not cross the  $p_{1/2}$  or  $f_{3/2}$ 

Osc. No.	Square well	Spect. term	Spin term	No. of states	Shells	Total No.
0	1.5	Is	151/2	2	2	2
			101/2	4)		
1	10	2∌	1.01/2	2	6	8
2	[1d	34	$1d_{4/2}$	6]		
	- 1		$1d_{1/2}$	4	12	
	25	25	251/2	2		20
3	(1f	4f	1fr/s	8	87	28?
			1/1/2	6)	01	201
	2p	3 p	2 / 2/2	4	22	
			201/2	2		
			1 g 9/2	10		50
4	[1g	5g				
			1 gr/2	8		
	2d	48	2ds/2	6		
			$2d_{1/2}$	4	32	
	35	3s	351/2	2		
			1811/2	12]		82
5	1.4	6k				
			$1h_{9/2}$	10)		
	2 3p	5/	2f1/2	8		
			$2f_{8/2}$	6	44	
		40	3 \$2/2	4		
			3 \$ 1/2	2		
6			1/11/1	14		126
	16	71				
			1/11/8			
	2g	6g				
	3d	54				

On Closed Shells in Nuclei. II

MARIA GOEPPERT MAYER ional Laboratory and Department of Physics iversity of Chicago, Chicago, Illinois February 4, 1949

THE spins and magnetic moments of the even-odd nuclei have been used by Feenbergh' and Nordheim' to deter-mine the angular momentum of the eigenfunction of the odd particle. The tabulations given by them indicate that spin orbit coupling favors the state of higher total angular momentum. If strong spin-orbit coupling, increasing with angular momentum, is assumed, a level assignment different from either Feenberg or Nordheim is obtained. This assignment encounters a very few contradictions with experimental facts and requires no major crossing of the levels from those of a square well potential. The magic numbers 50, 82, and 126 occur at the place of the spin-orbit splitting of levels of high

occur at the place of the spin-orbit spitting of levels of mga angular momentum.

Table I contains in column two, in order of decreasing binding energy, the levels of the square well potential. The quantum number gives the number of radial nodes. Two levels of the same quantum number cannot cross for any type of potential well, except due to spin-orbit splitting. No evidence of any crossing is found. Column three contains the usual of any crossing is round. Column three contains the usual spectroscopic designation of the levels, as used by Nordheim and Feenberg. Column one groups together those levels which are degenerate for a three-dimensional isotropic oscillator potential. A well with rounded corners will have a behavior in between these two potentials. The shell grouping is given in column five, with the numbers of particles per shell and the total number of particles are and including each shell in column six and seven, respectively.

Within each shell he levels may be expected to be close in energy, and not necessarily in the order of the table, although the order of levels of the same pointly and the content of the column six and seven. The content of t

the order of levels of the same orbital angular momentum and different spin should be maintained. Two exceptions, 11Na<sup>23</sup>

or d levels should occur in this shell and there is no evidence

for any.

The only exception to the proposed assignment in this shell is the spin 5/2 instead of 7/2 for Mn<sup>48</sup>, and the fact that

indicates  $f_{5/2}$  instead of the predicted  $d_{5/2}$ . No  $h_{11/2}$  levels appear. It seems that these levels are filled in pairs only,

LETTERS TO THE EDITOR

which does not seem a serious drawback of the theory as this tendency already shows up at the filling of the  $g_{\theta/t}$  levels. Otherwise, the agreement is satisfactory. The shell begins with  $_{45}$ Sb, which has two isotopes with  $d_{4/2}$  and  $g_{7/2}$  levels, respectively, as it should. The thallium isotopes with 81 neutrons and a spin of 1/2 indicate a crossing of the  $h_{11/2}$  and 3s levels. This is not surprising, since the energies of these levels are close together in the square well. The assignment

have been found. No f or p levels should occur and, except for

and the data indicates none The prevalence of isomerism towards the end of a shell, noticed by Feenberg and Nordheim, is easily understood by this assignment. These are the regions where levels with very different spins are adjacent. These ground and isomeric

Thanks are due to Enrico Fermi for the remark "Is there

state, is a heautiful confirmation of the correct heginning of

the next shell. Here information begins to be scarce. The spin and magnetic moment of Pb<sup>tot</sup> with 125 neutrons interpret as  $p_{1/5}$ . This is the expected end of the shell since 7s and 4p have practically the same energy in the square well model. No spins of 11/2 and no s, d, or g orbits should occur in this shell,



Enrico Fermi (1901–1954)

Thanks are due to Enrico Fermi for the remark, "Is there any indication of spin-orbit coupling?" which was the origin of this paper.

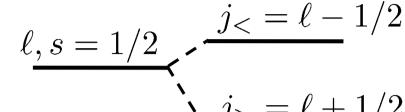
#### Shell (single particle) structures in nuclei

#### On Closed Shells in Nuclei, II

#### MARIA GOEPPERT MAYER

Argonne National Laboratory and Department of Physics, University of Chicago, Chicago, Illinois

February 4, 1949



free are the only violations. own spins and orbital assignments from ten these are known and unambiguous,

of the expected  $d_{5/2}$ , and  $d_{25}$  Mn<sup>25</sup> with 5/2

LETTERS TO THE EDITOR

which does not seem a serious drawback of the theory as this tendency already shows up at the filling of the  $g_{\theta/t}$  levels. Otherwise, the agreement is satisfactory. The shell begins with  $_{45}$ Sb, which has two isotopes with  $d_{4/2}$  and  $g_{7/2}$  levels, respectively, as it should. The thallium isotopes with 81 neutrons and a spin of 1/2 indicate a crossing of the  $k_{11/2}$  and 3s levels. This is not surprising, since the energies of these levels are close together in the square well. The assignment

TABLE II. Spins of even-odd nuclei.

Odd neutron O=neutron

state, is a beautiful confirmation of the correct beginning of the next shell. Here information begins to be scarce. The spin and magnetic moment of Pb<sup>tor</sup> with 125 neutrons interpret as and magnetic moment of  $P^{(i)}$  with 123 neutrons interpret as  $p_{1/2}$ . This is the expected end of the shell since 7i and 4p have practically the same energy in the square well model. No spins of 11/2 and no s, d, or g orbits should occur in this shell, and the data indicates none. The prevalence of isomerism towards the end of a shell, noticed by Feenberg and Nordheim, is easily understood by

The spin and magnetic moment of siBi, indicating an  $h_{9/2}$ 

this assignment. These are the regions where levels with very different spins are adjacent. These ground and isomeric states should also have different parity. Thanks are due to Enrico Fermi for the remark, "Is there

y indication of spin-orbit coupling?" which was the origin



actual place of the spin-out spiriting of reversion in angular momentum.

Table I contains in column two, in order of decreasing binding energy, the levels of the square well potential. The quantum number gives the number of radial nodes. Two levels of the same quantum number cannot cross for any type of potential well, except due to spin-orbit splitting. No evidence of any crossing is found. Column three contains the usual spectroscopic designation of the levels, as used by Nordheim and Feenberg. Column one groups together those levels which are degenerate for a three-dimensional isotropic oscillator potential. A well with rounded corners will have a behavior in between these two potentials. The shell grouping is given in column five, with the numbers of particles per shell and the total number of particles up to and including each shell in

levels, the first spin of 9/2 should occur at 41, which is indeed the case. Three nuclei with N or Z=49 have  $g_{97}$  orbits. No sor d levels should occur in this shell and there is no evidence

The only exception to the proposed assignment in this shell is the spin 5/2 instead of 7/2 for Mn<sup>55</sup>, and the fact that the magnetic moment of <sub>27</sub>Co<sup>59</sup> indicates a g<sub>7/2</sub> orbit instead of the expected  $f_{7/9}$ .

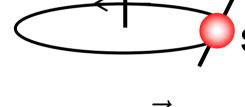
column six and seven, respectively.

Column six and seven, respectively.

Within each shelf the levels may be expected to be close in sergy, and not necessarily in the order of the table, although the order of levels of the same orbital angular momentum and different spin should be maintained. Two exceptions,  $\ln N^{2\alpha}$  adiquates  $f_{ij}$  instead of the predicted  $d_{dis}$ . No  $\theta_{ij1}$  levels different spin should be maintained. Two exceptions,  $\ln N^{2\alpha}$  adiquates  $f_{ij1}$  instead of the predicted  $d_{dis}$ . No  $\theta_{ij1}$  levels and different spin should be maintained. Two exceptions,  $\ln N^{2\alpha}$  and  $d_{ij2}$  instead of the predicted  $d_{dis}$ . No  $\theta_{ij1}$  levels are the predicted  $d_{dis}$ . No  $\theta_{ij1}$  levels are the predicted  $d_{dis}$  in  $d_{ij2}$  and  $d_{ij3}$  and  $d_{ij4}$  are the predicted  $d_{ij4}$ . No  $\theta_{ij1}$  levels are the predicted  $d_{ij4}$ . No  $\theta_{ij1}$  levels are the predicted  $d_{ij4}$  in  $d_{ij4}$  and  $d_{ij4}$  are the predicted  $d_{ij4}$ . No  $\theta_{ij4}$  levels  $d_{ij4}$  are the predicted  $d_{ij4}$  in  $d_{ij4}$  and  $d_{ij4}$  are the predicted  $d_{ij4}$ . No  $\theta_{ij4}$  levels  $d_{ij4}$  and  $d_{ij4}$  are the predicted  $d_{ij4}$  in  $d_{ij4}$  and  $d_{ij4}$  are the predicted  $d_{ij4}$  in  $d_{ij4}$  and  $d_{ij4}$  are the predicted  $d_{ij4}$  in  $d_{ij4}$  and  $d_{ij4}$  are the predicted  $d_{ij4}$  are the predicted  $d_{ij4}$  and  $d_{ij4}$  are the predicted  $d_{ij4}$  are the predicted  $d_{ij4}$  and  $d_{ij4}$  are the predicted  $d_{ij4}$  and  $d_{ij4}$  are the predicted  $d_{ij4}$  and  $d_{ij4}$  are the predicted  $d_{ij4}$  are the predicted  $d_{ij4}$  and  $d_{ij4}$  are the predicted  $d_{ij4}$  and

B		Odd prosons		Out neutron			X = proton				# #		
No. of neutrons or protons	Element	Mass No. orbit	Mans No. orbit	Element	Mass No. orbit	1/2	î i	spin d 2 5/2	<i>f</i> 3	9/2	No. odd 20 or	Levels	
1	Н	1 #	3 #	He	3 s	8					1	1#	
5	Li B N	7 p 11 p 15 p		Be C	9 p 13 p	8	X ⊗				3 5 7	P1/2 P1/2	
11 13 15	F Na Al P Cl K	19 # 23 27 d 31 35 d 39 d	37 d 41 d	s	33	X	x R	X			9 11 13 15 17	d <sub>1/2</sub> 81/2 d <sub>1/2</sub>	
23 25	Se V Mn Co	45 f 51 55 59 g						х	X X		21 23 25 27	fin	
31 33 35 37 39 41 43 45	Cu Ga As Br Rb Y Cb Te Rh	63 69 75 79 85 f(5/2) 89 93 g	65 81 87 p(3/2)	Zn Se	67 f 77	<i>x</i> <i>o</i>	X X X X	8		x	29 31 33 35 37 39 41 43 45	Pive five Pive Ove	
47 49	Ag In	107 p 113 g	109 p 115 g	Kr Sr	83 g 87 g	X				8	47 49		
55 57 59	Sb I Cs La Pr Pm	$^{121}_{127}  ^{d(5/2)}_{133  g}_{139  y}$	123 g(7/2) 137 g	Мо	95	0		X X	X X		51 53 55 57 59 61	g://2 di/2	
63 65 67 69 71	Eu Tb Ho Tm Lu	151 159 165	153 f	Cd Cd Sn Sn	111 s s, Sa s 117 s 119 s	0000	х	х	X X		63 65 67 69	$k_{11/2}$	
75 77 79	Ta Re Ir Au Tl	181 g 185 191 (1/2) 197 d 203	187 193 (3/2) 205	Xe Xe Ba Ba	129 s 131 d 135 d 137 d	o X	880	X	A		71 73 75 77 79 81	dap2 81/2	
83	Bi	209 h				_				X	83	h1/2	

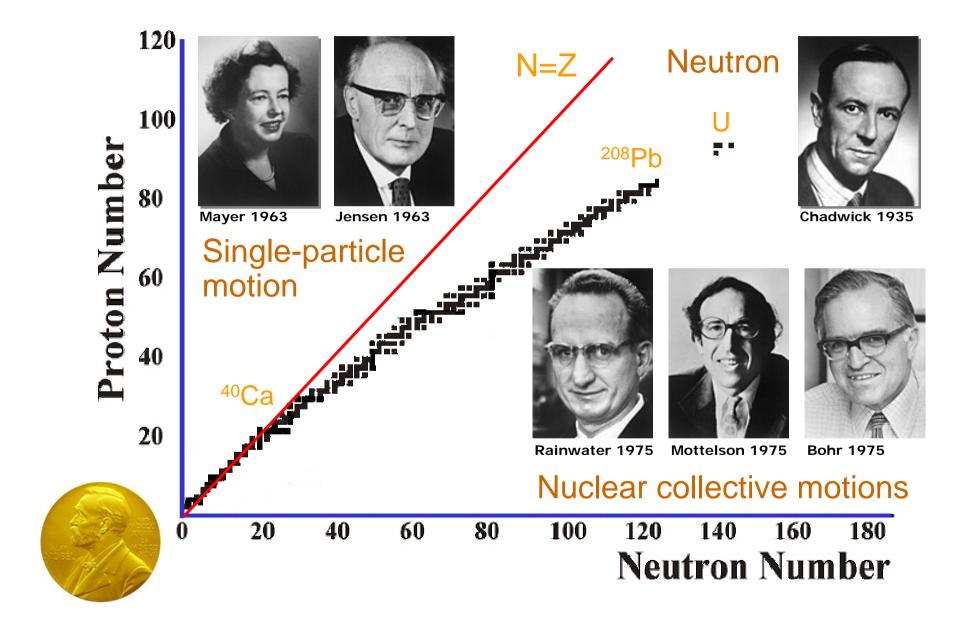
demands that there be no spins of 9/2 in this shell, and none have been found. No f or p levels should occur and, except for Euis, there is no indication of any



 $V_{\ell s}(r)\vec{\ell}\cdot\vec{s}$ 

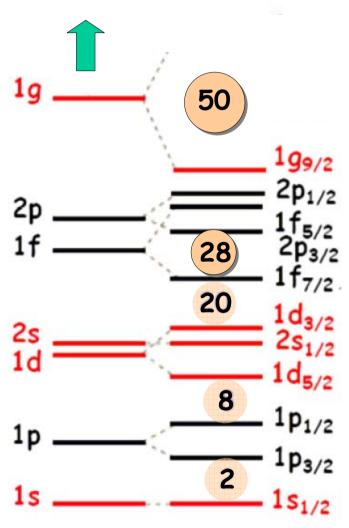
Thanks are due to Enrico Fermi for the remark, "Is there any indication of spin-orbit coupling?" which was the origin of this paper.

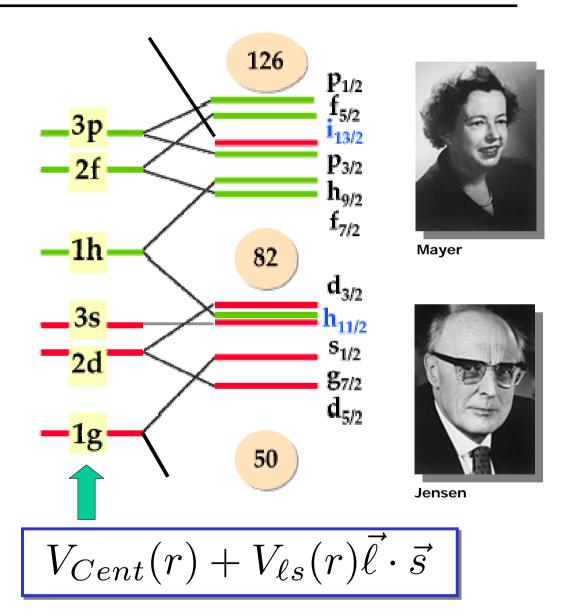
#### 300 stable isotopes exist on Earth – Segré chart



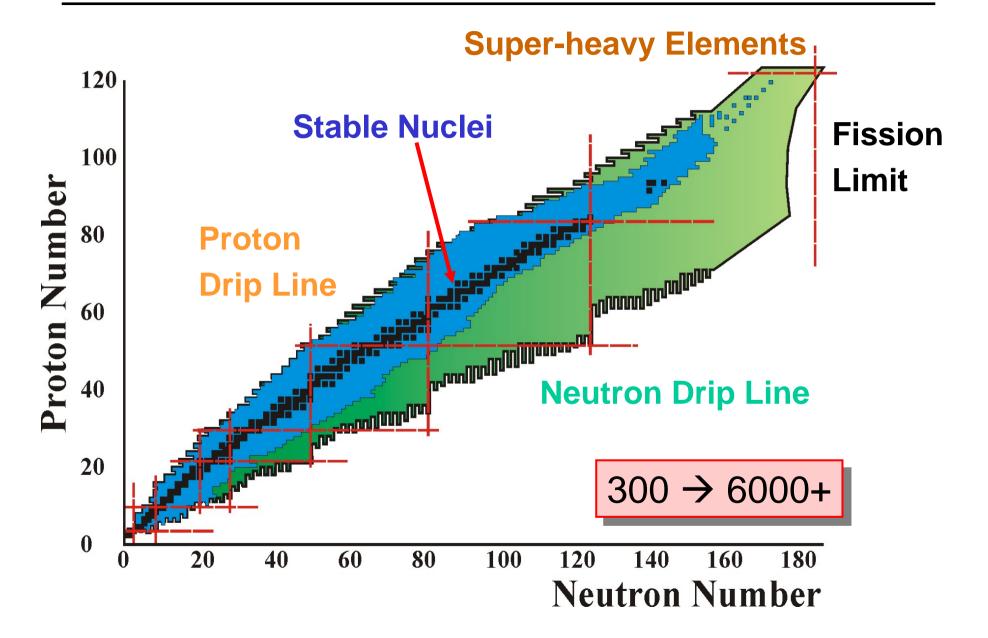
#### Magic numbers – a framework for nuclear stability



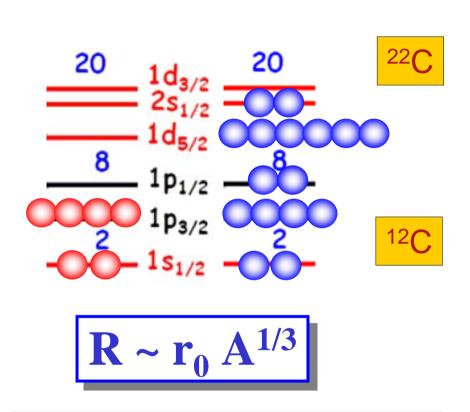




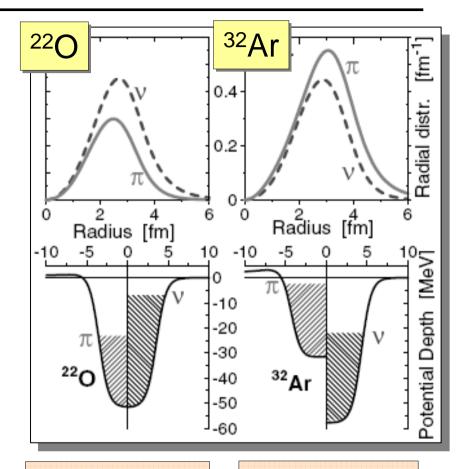
#### So, the limits to nuclear existence are ...?



#### Exotic neutron + proton combinations - expectations



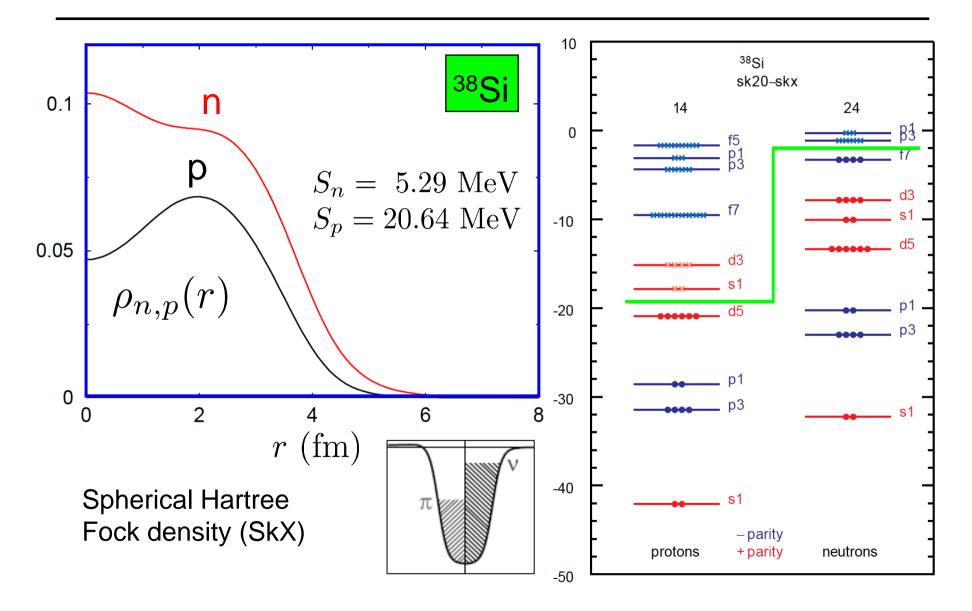
Very weak binding of last nucleons as the <u>driplines</u> are approached (1.4 MeV in <sup>22</sup>C, 18.7 MeV for neutron in <sup>12</sup>C)



Z=8 **N=14**  $S_n=6.8$  MeV  $S_p=23$  MeV

Z=18 **N=14**  $S_n=22$  MeV  $S_p=2.4$  MeV

#### Make use of available tools – locate Fermi surfaces



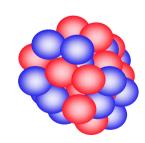
#### Bound states – the mean field helps our intuition

```
INPUT VALUES
                  IA,IZ =
  ---- Neutron bound state results
                       OCC
 k n l
                   IE
 1 1 s 1/2 -26.757 1
                       2.00
                              36.70
                                     35.28
 2 1 p 3/2 -16.883 1
                       4.00
                              36.70
                                     35.80
  3 1 p 1/2 -12.<u>396 1</u>
                              36.70
                       2.00
                                     36.04
 4 1 d 5/2
            -6.166 1
                       6.00
                              36.70
                                     36.37
  5 1 d 3/2
             -0.109 1
                       0.00
                              36.70
                                     36.69
           -3.360 1
                       2.00
                             36.70
                                     36.52
                       0.00
           -0.200 3
                              46.02
                                     46.01
  8 1 f 5/2
            -0.200 3
                       0.00
                             60.56
                                     60.55
  9 2 p 3/2
            -0.2003
                       0.00 48.10
                                     48.09
 ---- Neutron single-particle radii -----
```

The mean field – e.g. spherical HF - gives an excellent estimate and guide

$$\langle r^2 \rangle = \frac{A}{A-1} \langle r^2 \rangle_{HF}$$

```
R(2)
                   R(4)
                          OCC
                                  rho(8.9) rho(9.9) rho(10.9)
1.1 s 1/2
           2.274
                  2.575
                          2.000
                                 0.848E-09 0.706E-10 0.600E-11
2 1 p 3/2
           2.863
                  3.133
                         4.000
                                 0.188E-07 0.244E-08 0.325E-09
3 1 p 1/2
           2.954
                  3.268
                         2.000
                                 0.727E-07 0.122E-07 0.210E-08
4 1 d 5/2
           3.434
                  3.757
                         6.000
                                 0.524E-06 0.129E-06 0.327E-07
           4.662
                  6.063
                         0.000
                                 0.131E-04 0.675E-05 0.371E-05
6 2 s 1/2
           4.172
                  4.895
                         2.000
                                 0.769E-05 0.278E-05 0.102E-05
           3.865 4.440
                         0.000
                                 0.324E-05 0.134E-05 0.600E-06
8 1 f 5/2
           3.890
                  4.477
                         0.000
                                 0.341E-05 0.141E-05 0.631E-06
9 2 p 3/2
           6.815
                  8.635
                         0.000
                                 0.451E-04 0.270E-04 0.167E-04
```



$$^{24}O(g.s.)$$

#### And more realistic (complex) structure models?

MARIA GOEPPERT MAYER Nobel Lecture, December 12, 1963

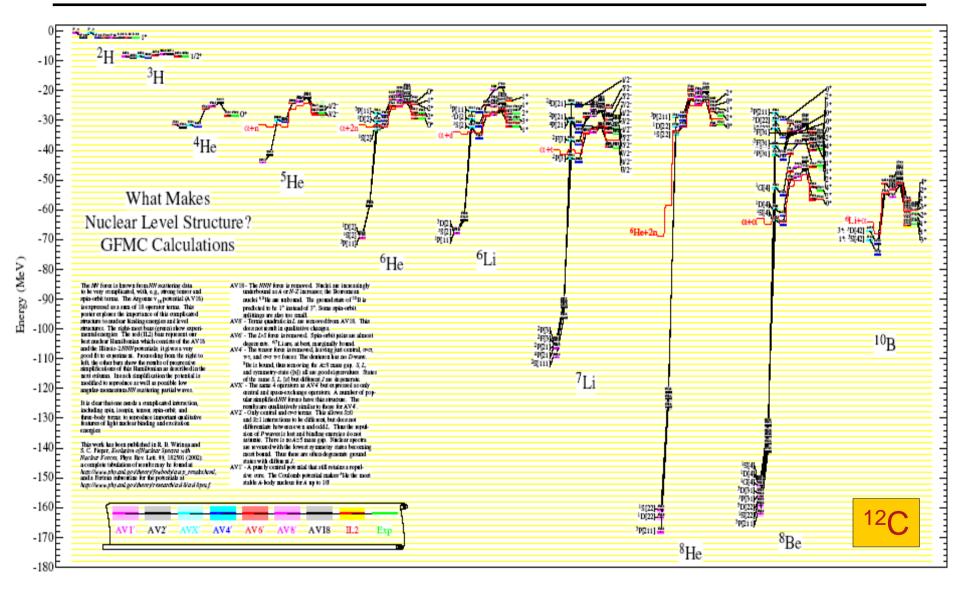
The shell model → more properly: <u>independent particle model</u> .....

If the forces are known, one should in principle be able to calculate deductively the properties of individual complex nuclei. Only after this has been accomplished can one say that one completely understands nuclear structures.

Ground state of <sup>12</sup>C 75,000 CPU-hours (27 Teraflop hours)

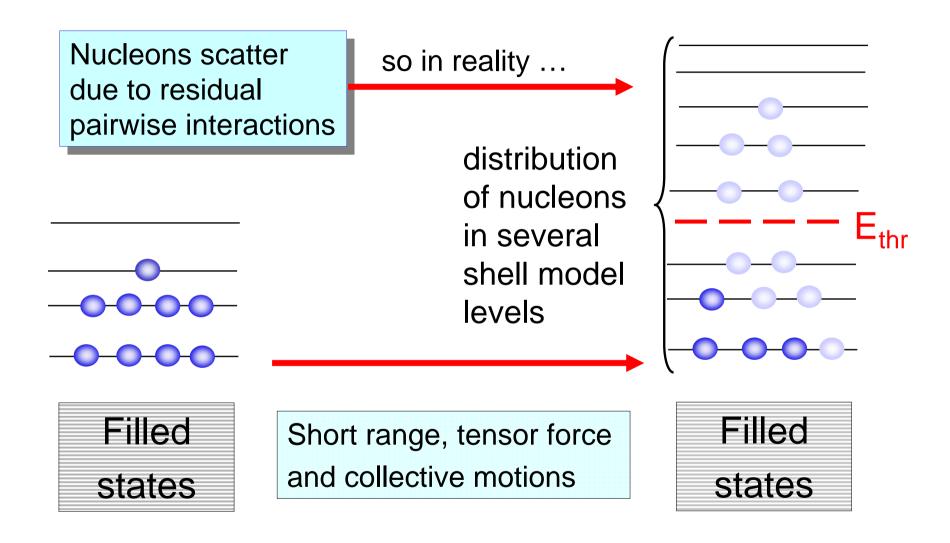
The shell model has initiated a large field of research. It has served as the starting point for more refined calculations. There are enough nuclei to investigate so that the <u>shell modellists will not soon be unemployed.</u>

#### The state-of-the art – A=12 and less – really hard!

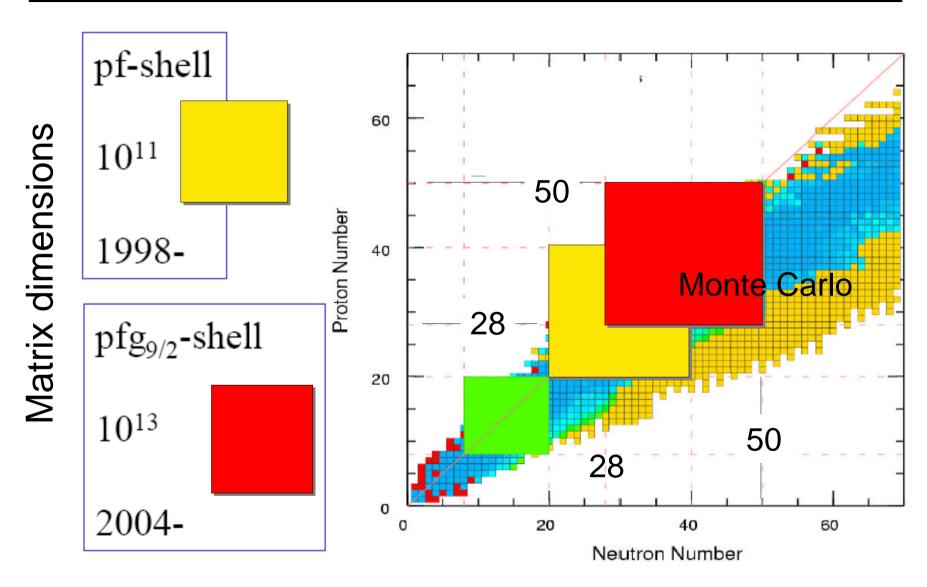


http://www.phy.anl.gov/theory/poster.pdf

#### Modern 'shell model' calculations do much more ....

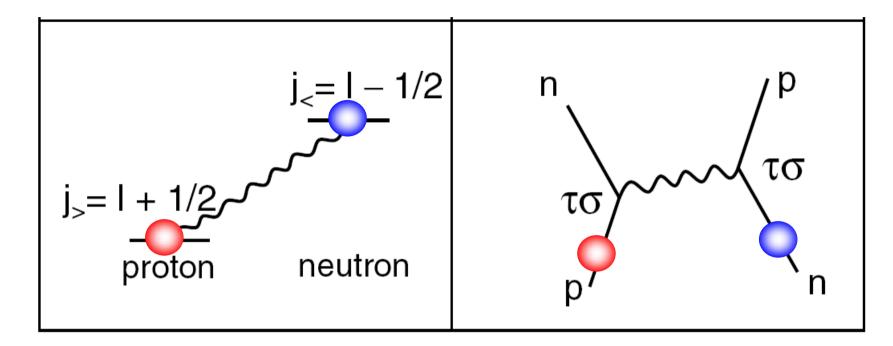


### Calculations are big - computational challenge



From Alex Brown

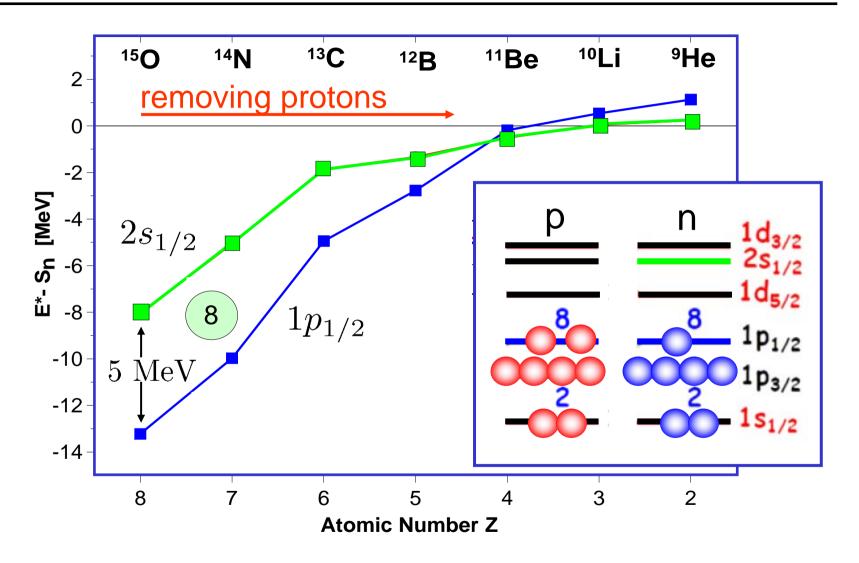
#### Implications for shell structure: asymmetric nuclei



Attractive interaction between neutrons and protons occupying  $j_{>}$  and  $j_{<}$  levels, <u>repulsive</u>  $j_{>}$  and  $j_{>}$  levels – from several sources, but mainly the tensor force

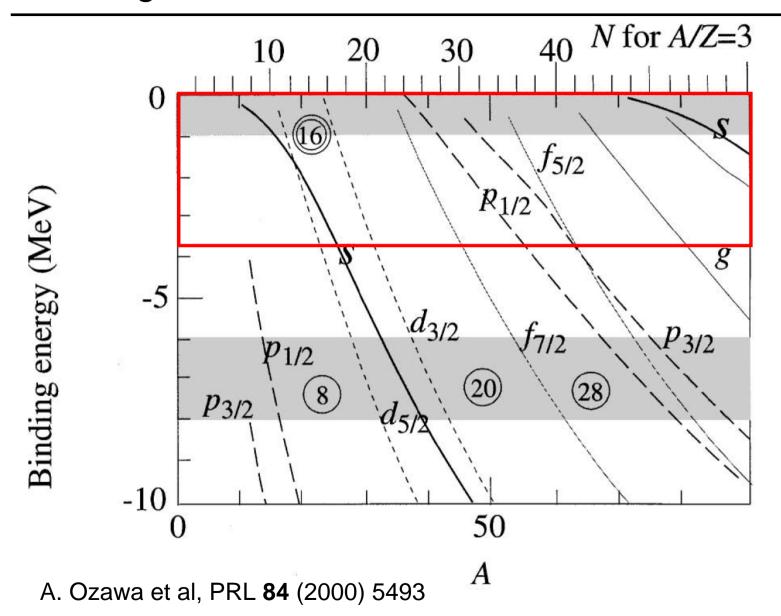
Takaharu Otsuka et al, Phys. Rev. Lett. **87**, 082502 (2001), **95**, 232502 (2005), **105**, 032501 (2010)

#### Migration of levels for N=7 – loss of the N=8 magic #

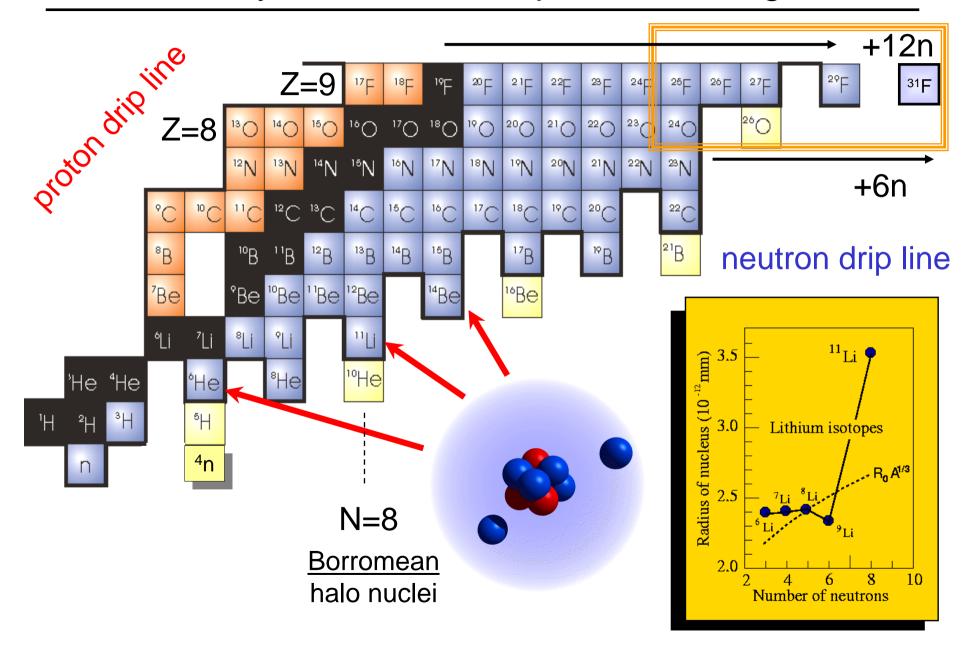


From: P.G. Hansen and J.A. Tostevin, Ann Rev Nucl Part Sci 53 (2003) 219

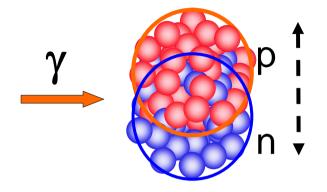
#### Low angular momentum states see well diffuseness



#### We have only reached the driplines in the light nuclei

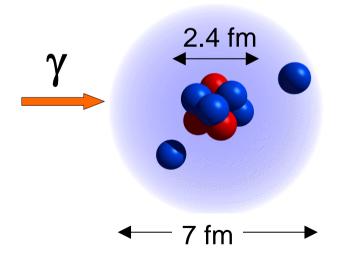


#### Halo nuclei –extremes of weak binding - responses

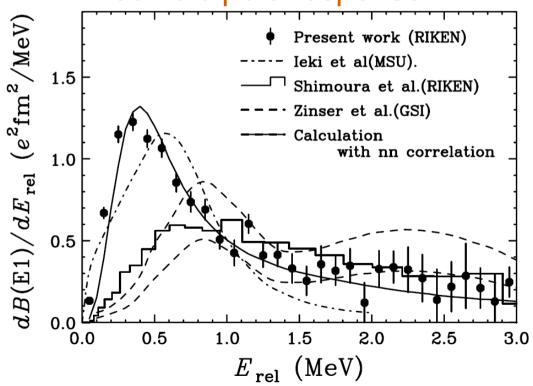


<sup>6</sup>He (2n,1 MeV) <sup>11</sup>Be (1n,0.5 MeV) <sup>11</sup>Li (2n,0.5 MeV) <sup>14</sup>Be (2n,~1 MeV)

(Giant) electric Dipole excitation → 10-20 MeV

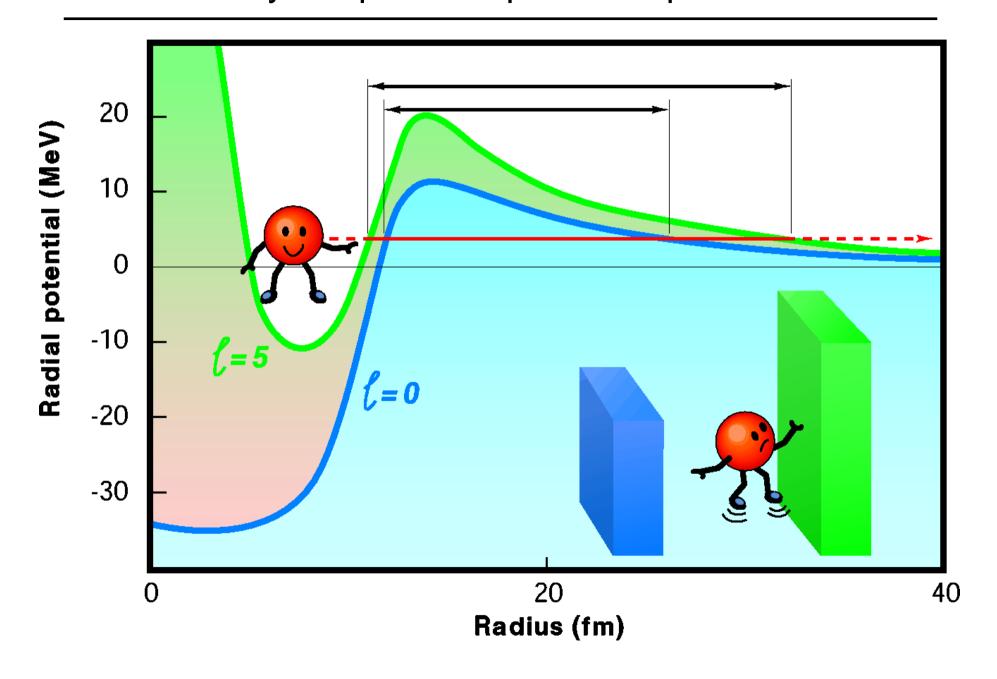


#### Electric dipole response – <sup>11</sup>Li

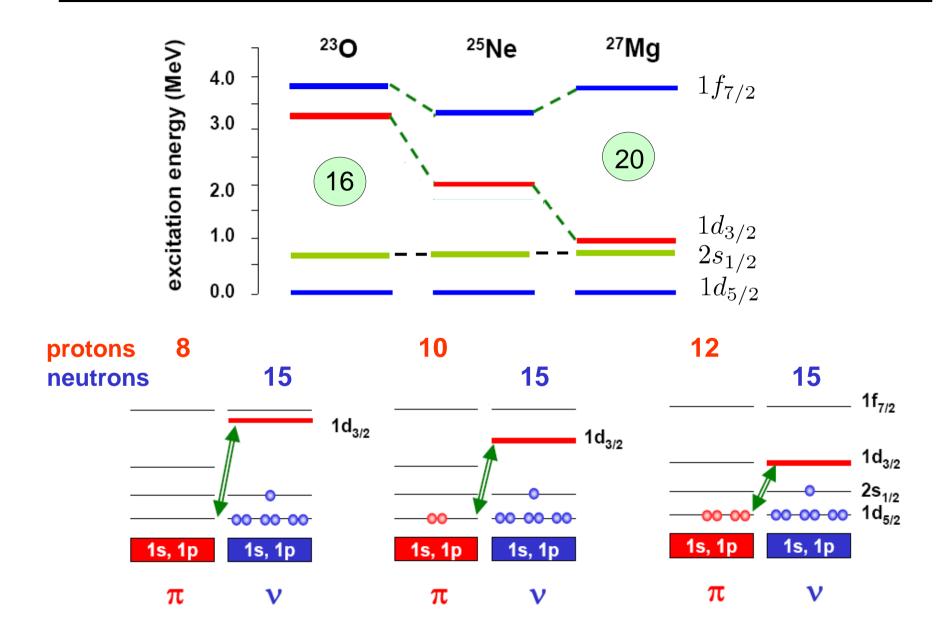


T. Nakamura et al., PRL (2006) in press

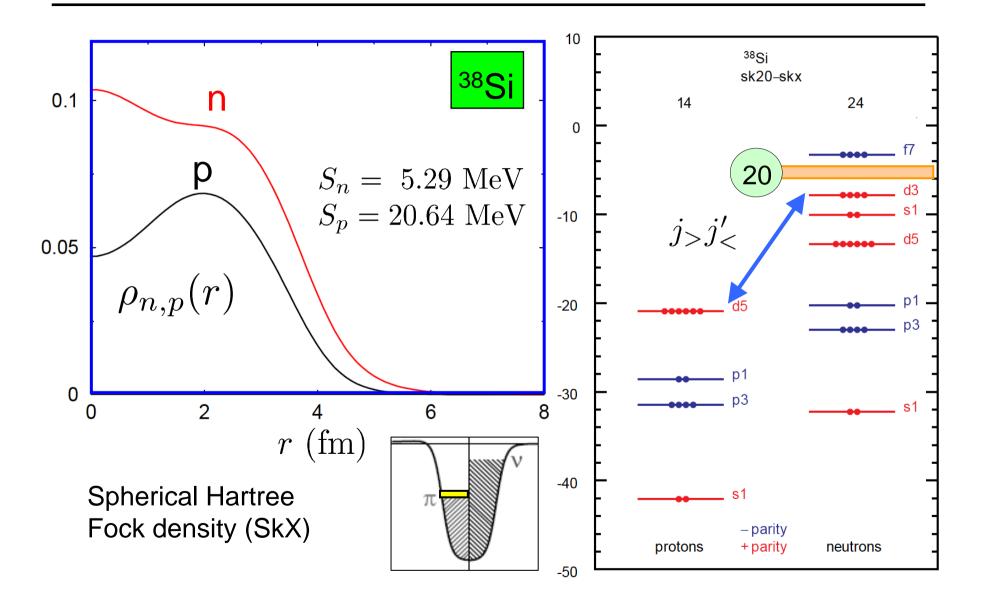
#### Proton decay – a probe of proton-dripline structures



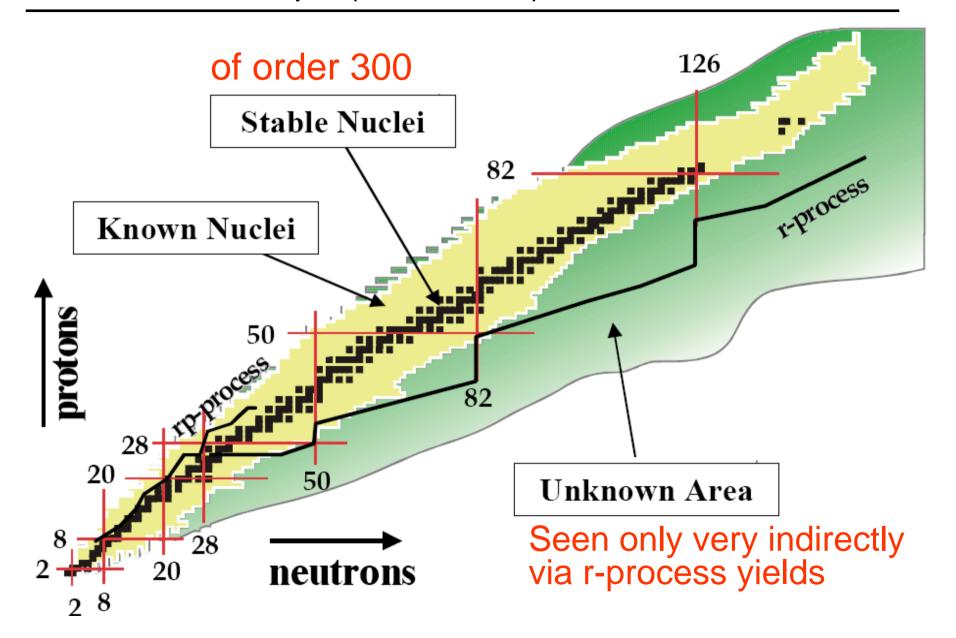
# Magic numbers change with "neutron richness"



#### Otsuka: the np interaction tensor correlation

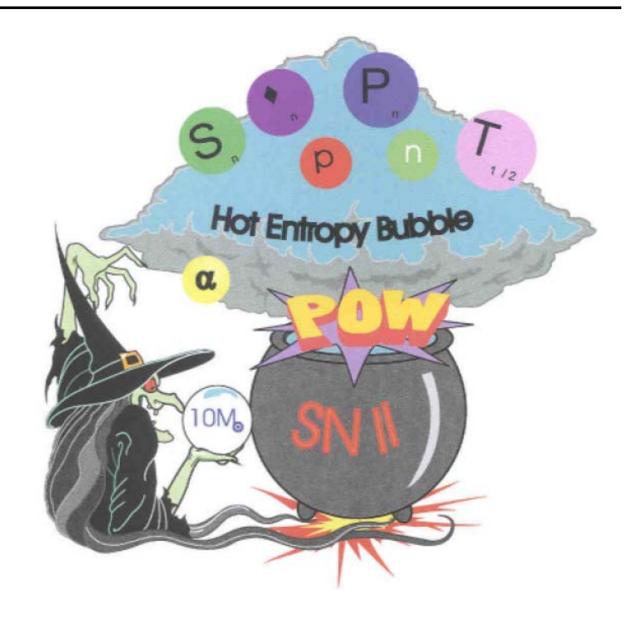


#### Nuclear landscape (circa 2010)



# ½ of all --- is produced in r-process synthesis

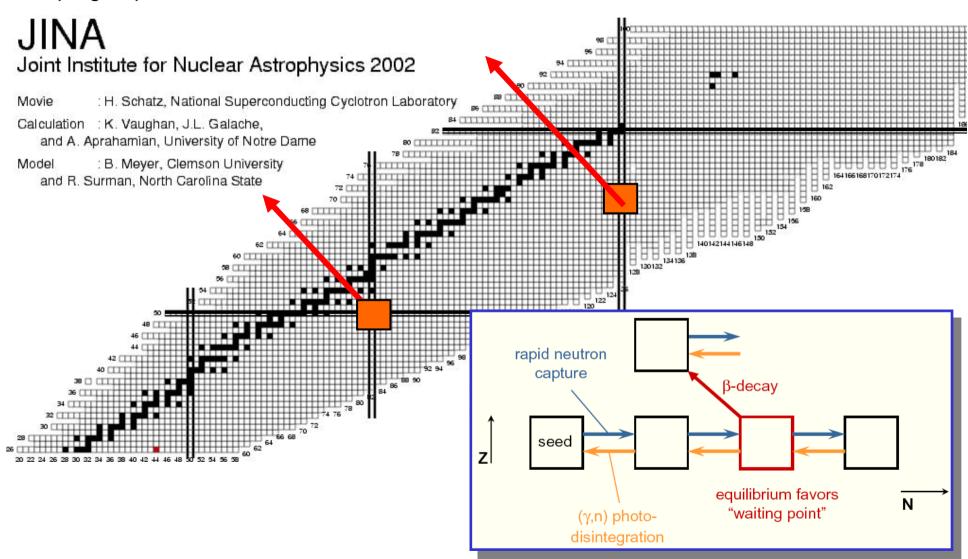




#### The Joint Institute for Nuclear Astrophysics



http://groups.nscl.msu.edu/nero/Web/materials.html

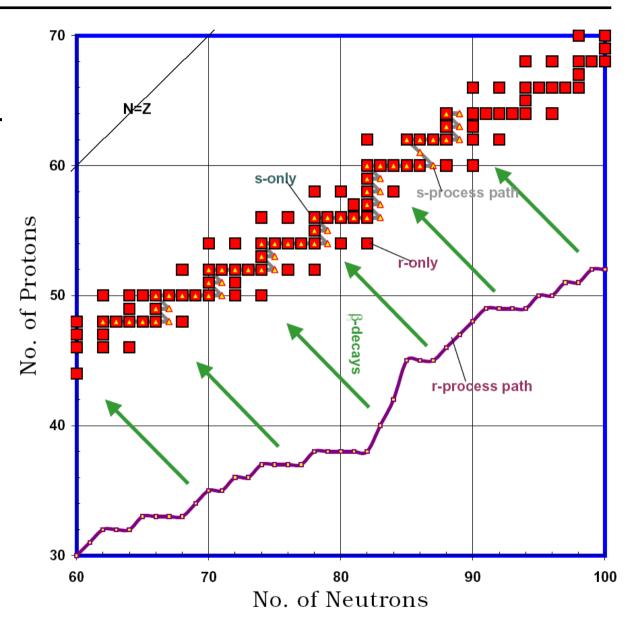


### r-process synthesis modelling – inputs needed?

#### Need is for:

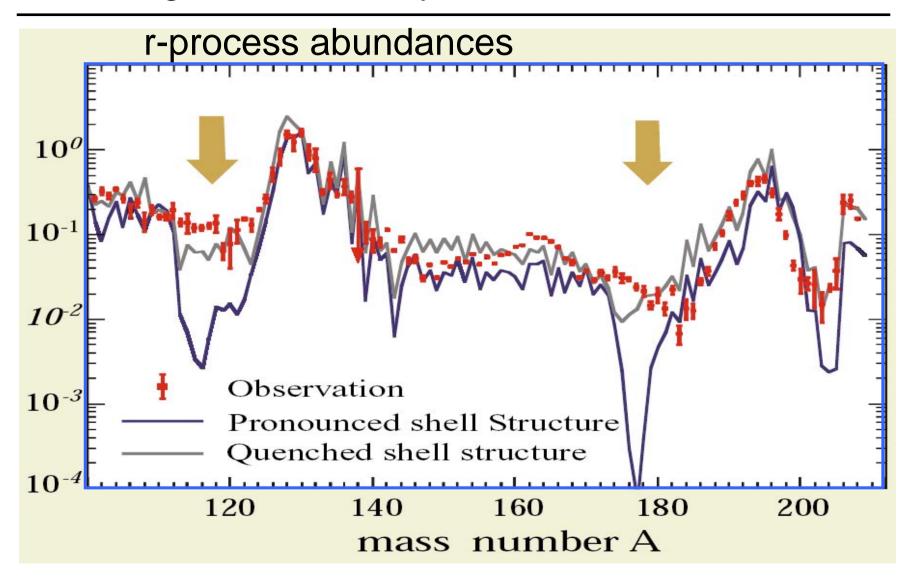
Masses – stability - determine reaction rates

beta-decay halflives – masses, but also structure effects – need to test shell model predictions – too many systems to measure sensibly



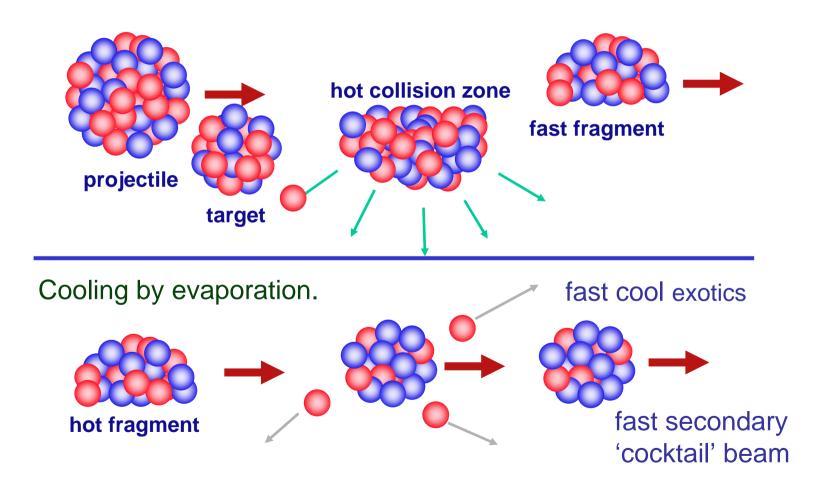


# 'Less magic' in heavier systems – evidence?



### Exotic nuclei production - projectile fragmentation

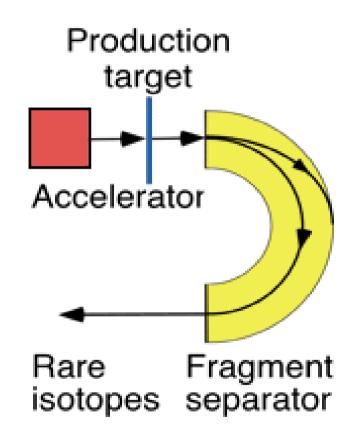
Random removal of protons and neutrons from heavy projectile in peripheral collisions at <u>high energy</u> - 100 MeV per nucleon or more



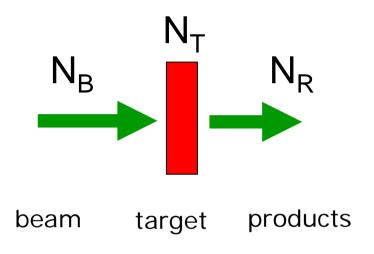
# Schematic of a Projectile Fragmentation Facility

GANIL (France), GSI (Germany), NSCL (USA), RIKEN (Japan)

- High-energy beams (E/A > 50 MeV) of only modest beam quality
- Fast, physical method of separation, no chemistry
- Suitable for short-lived isotopes (lifetimes > 10<sup>-6</sup> s)
- But, low-energy beams are (very) difficult
- Modest beam quality



#### Use of rarest beams a 'few' atomic nuclei/second



- Fast exotic beams allow for
  - thick secondary targets
  - event-by-event identification
  - clean product selection
  - nevertheless .....

- $N_R = s \times N_T \times N_B$ 
  - s cross section
  - N<sub>T</sub> atoms in target
  - N<sub>B</sub> beam rate
  - N<sub>R</sub> reaction rate

- Example
  - s = 100 millibarn

$$N_T = 10^{21}$$

$$N_B = 3 Hz$$

• 
$$N_R = 26/day$$
  
=  $3 \times 10^{-4} Hz$ 

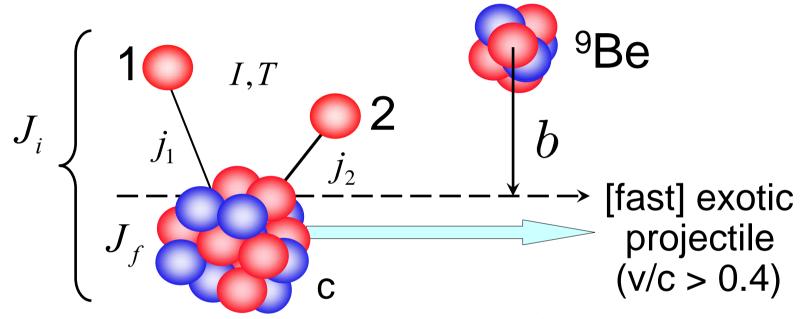
From Thomas Glasmacher

#### So, learned so far

- 1: Must rethink 'textbook' nuclear shell structure when treating asymmetric N:Z systems see breakdown of N=8, 20, (28 ...?) shell gaps.
- 2: Asymmetry allows the observation and study of novel NN interaction effects and new structure
- 3: An effective means of production is using high energy fragmentation that produce the exotic nuclei as fast secondary beams: 100 MeV/u
- 4. Experiments for the rarest cases, with intensities of a few particles/s are hard reactions with high cross sections/efficient detection are essential
- 5: Structure calculations for A>12 are hard, they use effective interactions that need to be tested

# Probing single particle (shell model) states

One such experimental option is one or two-nucleon removal – at ~100 MeV/nucleon

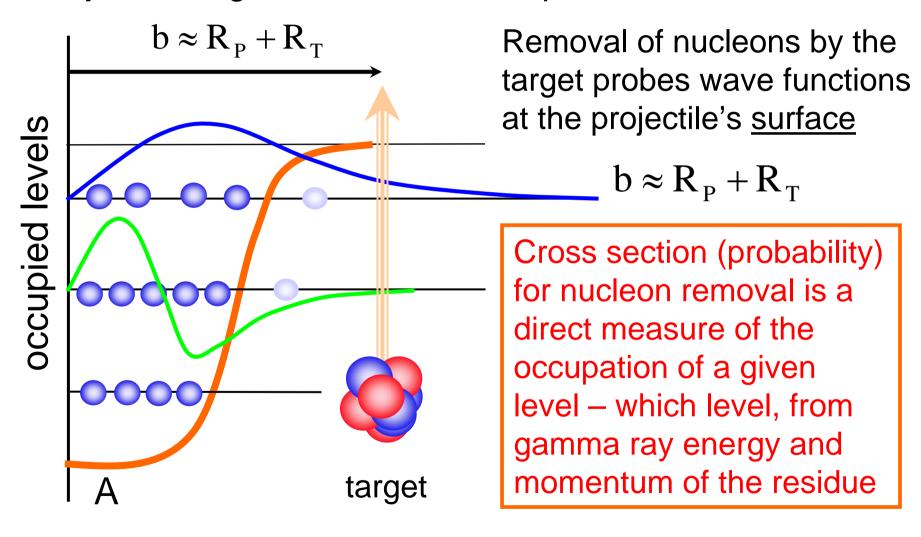


Experiments do not measure target final states. Final state of core c measured – using decay gamma rays.

How can we describe and what can we learn from these?

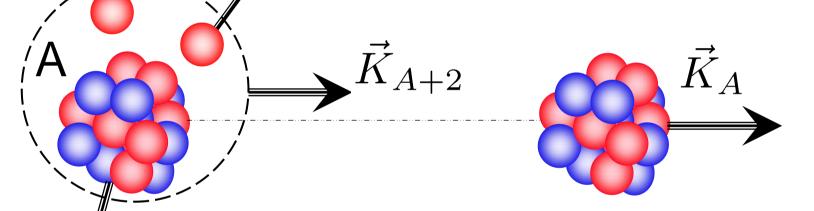
#### Viewed from the rest frame of the projectile

Projectile-target interaction absorptive when come close



#### Sudden 2N removal from the mass A residue

Sudden removal: residue momenta probe the summed momenta of pair in the projectile rest frame



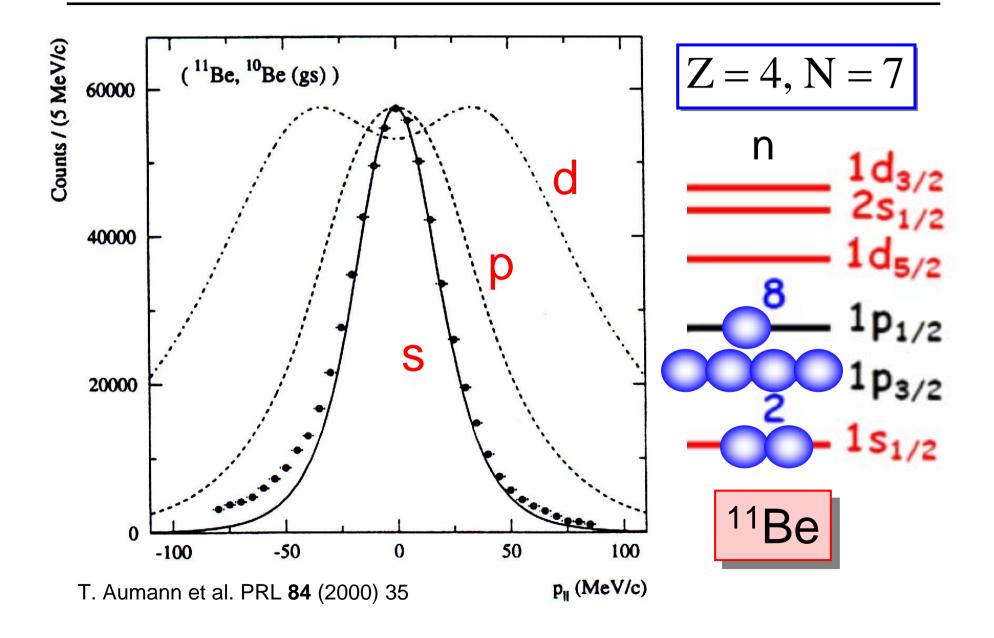
Projectile rest frame

$$\vec{K}_A = \frac{A}{A+2}\vec{K}_{A+2} - [\vec{\kappa}_1 + \vec{\kappa}_2]$$

laboratory frame  $\vec{K}_A$  and  $\vec{K}_{A+2}$ 

and component equations

#### Residue momentum <sup>11</sup>Be→ <sup>10</sup>Be – halo case



## Need description of high-energy nuclear collisions

Need composite particle (nucleons and core), but proceed in steps:

- 1. Point particle scattering summary only
  - S-matrix and the eikonal (high energy, forward-focussed reaction) approximation
- 2. Composite few-particle system
  - adiabatic/sudden approximation (reaction fast)
  - adiabatic plus eikonal scattering solution
- 3. Equations for one and two-nucleon removal cross sections and other observables
- 4. Link of reactions & many-body nuclear structure calculations (and assumptions made/needed)

#### Point particles: phase shift and partial wave S-matrix

#### Scattering states

$$E_{cm} > 0$$
  $k = \sqrt{\frac{2\mu E_{cm}}{\hbar^2}}$ 

$$\left(\frac{d^2}{dr^2} - \frac{\ell(\ell+1)}{r^2} - \frac{2\mu}{\hbar^2} U_{\ell j}(r) + k^2\right) u_{k\ell j}(r) = 0$$

and beyond the range of the nuclear forces, then

$$\left(\frac{d^2}{dr^2} - \frac{\ell(\ell+1)}{r^2} - \frac{2\eta k}{r} + k^2\right) u_{k\ell j}(r) = 0, \quad \eta = \frac{\mu Z_c Z_v e^2}{\hbar k}$$

 $F_{\ell}(\eta,kr),~G_{\ell}(\eta,kr)~$  regular and irregular Coulomb functions

$$u_{k\ell j}(r) \rightarrow e^{i\delta_{\ell j}} \left[\cos \delta_{\ell j} F_{\ell}(\eta, kr) + \sin \delta_{\ell j} G_{\ell}(\eta, kr)\right]$$

$$\rightarrow (i/2) \left[H_{\ell}^{(-)}(\eta, kr) - S_{\ell j} H_{\ell}^{(+)}(\eta, kr)\right]$$

$$H_{\ell}^{(\pm)}(\eta,kr) = G_{\ell}(\eta,kr) \pm iF_{\ell}(\eta,kr)$$

## S-matrix - ingoing and outgoing waves amplitudes

$$u_{k\ell j}(r) \to (i/2) [1] H_{\ell}^{(-)} - S_{\ell j} H_{\ell}^{(+)}]$$

$$E \longrightarrow 1 H_{\ell}^{(-)} S_{\ell j} H_{\ell}^{(+)} \Rightarrow$$

$$W(r) \longrightarrow r$$

$$|S_{\ell j}|^2$$

$$(1 - |S_{\ell j}|^2)$$

## Eikonal approximation: for point particles (1)

Approximate (semi-classical) scattering solution of

$$\left(-\frac{\hbar^2}{2\mu}\nabla_r^2 + U(r) - E_{cm}\right)\chi_{\vec{k}}^+(\vec{r}) = 0, \qquad \mu = \frac{m_c m_v}{m_c + m_v}$$
$$\left(\nabla_r^2 - \frac{2\mu}{\hbar^2}U(r) + k^2\right)\chi_{\vec{k}}^+(\vec{r}) = 0$$

valid when  $|U|/E \ll 1, \ ka \gg 1$   $\rightarrow$  high energy Key steps are: (1) the distorted wave function is written

$$\chi_{\vec{k}}^+(\vec{r}) = \exp(i\vec{k}\cdot\vec{r})$$
  $\omega(\vec{r})$  all effects due to  $U(r)$ , modulation function

(2) Substituting this product form in the Schrodinger Eq.

$$\left[2i\vec{k}\cdot\nabla\omega(\vec{r}) - \frac{2\mu}{\hbar^2}U(r)\omega(\vec{r}) + \nabla^2\omega(\vec{r})\right]\exp(i\vec{k}\cdot\vec{r}) = 0$$

# Eikonal approximation: point neutral particles (2)

$$\left[2i\vec{k}\cdot\nabla\omega(\vec{r}) - \frac{2\mu}{\hbar^2}U(r)\omega(\vec{r}) + \nabla^2\omega(\vec{r})\right]\exp(i\vec{k}\cdot\vec{r}) = 0$$

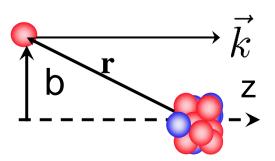
The conditions  $|U|/E \ll 1$ ,  $ka \gg 1 \rightarrow \text{imply that}$ 

$$2\vec{k} \cdot \nabla \omega(\vec{r}) \gg \nabla^2 \omega(\vec{r})$$

 $2\vec{k}\cdot\nabla\omega(\vec{r})\gg\nabla^2\omega(\vec{r})$  Slow spatial variation cf. k

and choosing the z-axis in the beam direction  $\vec{k}$ 

$$\frac{d\omega}{dz} \approx -\frac{i\mu}{\hbar^2 k} U(r)\omega(\vec{r})$$



$$\frac{d\omega}{dz} \approx -\frac{i\mu}{\hbar^2 k} U(r)\omega(\vec{r}) \qquad \text{phase that develops with z}$$
 with solution 
$$\omega(\vec{r}) = \exp\left[-\frac{i\mu}{\hbar^2 k} \int_{-\infty}^z U(r) dz'\right]$$

1D integral over a straight line path through U at the impact parameter b

## Eikonal approximation: point neutral particles (3)

$$\chi_{\vec{k}}^{+}(\vec{r}) = \exp(i\vec{k} \cdot \vec{r}) \ \omega(\vec{r}) \approx \exp(i\vec{k} \cdot \vec{r}) \ \exp\left[-\frac{i\mu}{\hbar^2 k} \int_{-\infty}^{z} U(r) dz'\right]$$

So, after the interaction and as  $z \rightarrow \infty$ 

$$\chi_{\vec{k}}^{+}(\vec{r}) \to \exp(i\vec{k} \cdot \vec{r}) \exp\left[-\frac{i\mu}{\hbar^2 k} \int_{-\infty}^{\infty} U(r) dz'\right] = S(b) \exp(i\vec{k} \cdot \vec{r})$$

$$\chi_{\vec{k}}^+(\vec{r}) \to S(b) \exp(i\vec{k} \cdot \vec{r})$$

Eikonal approximation to the S-matrix S(b)

$$S(b) = \exp\left[-\frac{i}{\hbar v} \int_{-\infty}^{\infty} U(r)dz'\right]$$

S(b) is amplitude of the forward going outgoing waves from the scattering at impact parameter b

$$v = \hbar k/m$$

Moreover, the structure of the theory generalises simply to few-body projectiles

# Eikonal approximation: point particles - summary

$$\chi_{\vec{k}}^{+}(\vec{r}) = \exp(i\vec{k} \cdot \vec{r}) \exp\left[-\frac{i\mu}{\hbar^{2}k} \int_{-\infty}^{z} U(r)dz'\right]$$

$$v = \hbar k/m$$

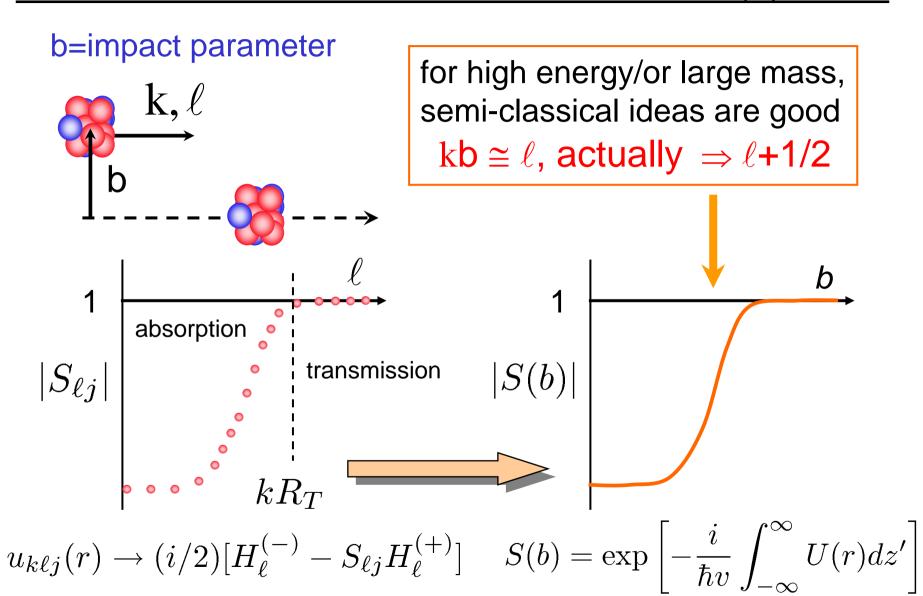
$$\chi(b) = -\frac{1}{\hbar v} \int_{-\infty}^{\infty} U(r)dz$$

$$U(r)$$

$$Z$$
limit of range of finite ranged potential
$$\chi_{\vec{k}}^{+}(\vec{r}) \to S(b) \exp(i\vec{k} \cdot \vec{r})$$

$$S(b) = \exp[i\chi(b)] = \exp\left[-\frac{i}{\hbar v} \int_{-\infty}^{\infty} U(r)dz'\right]$$

## Semi-classical models for the S-matrix - S(b)



#### Point particle scattering – cross sections

All cross sections, etc. can be computed from the S-matrix, in either the partial wave or the eikonal (impact parameter) representation, for example (spinless case):

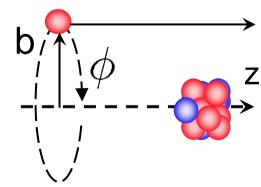
$$\sigma_{el} = \frac{\pi}{k^2} \sum_{\ell=0}^{\infty} (2\ell+1)|1 - S_{\ell}|^2 \approx \int d^2\vec{b} |1 - S(b)|^2$$

$$\sigma_R = \frac{\pi}{k^2} \sum_{\ell=0}^{\infty} (2\ell+1)(1 - |S_{\ell}|^2) \approx \int d^2\vec{b} (1 - |S(b)|^2)$$

$$\sigma_{tot} = \sigma_{el} + \sigma_R = 2 \int d^2 \vec{b} \left[ 1 - \mathrm{Re.} S(b) 
ight]$$
 etc.

and where (cylindrical coordinates)

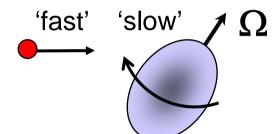
and where (cylindrical coordinates) 
$$\int d^2\vec{b} \equiv \int_0^\infty bdb \int_0^{2\pi} d\phi = 2\pi \int_0^\infty bdb$$



### Adiabatic (sudden) approximations in physics

Identify high energy/fast and low energy/slow degrees of freedom

Fast neutron scattering from a rotational nucleus

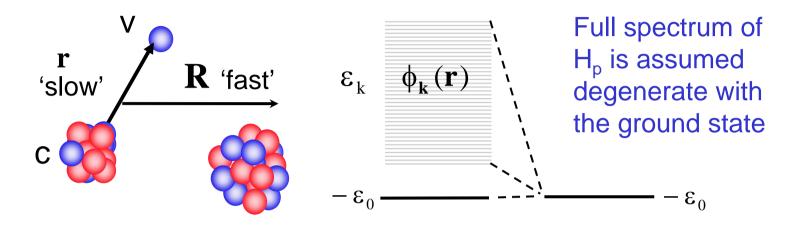


Fix  $\Omega$ , calculate scattering amplitude  $f(\theta, \Omega)$  for each (fixed)  $\Omega$ .

moment of inertia  $\rightarrow \infty$  and rotational spectrum is assumed degenerate

Transition amplitudes  $f_{\alpha\beta}(\theta) = \langle \beta \mid f(\theta, \Omega) \mid \alpha \rangle_{\Omega}$ 

#### Few-body projectiles – the adiabatic model



Freeze internal co-ordinate  $\mathbf{r}$  then scatter  $\mathbf{c}+\mathbf{v}$  from target and compute  $\mathbf{f}(\theta,\mathbf{r})$  for all required <u>fixed</u> values of  $\mathbf{r}$ 

Physical amplitude for breakup to state  $\phi_{\mathbf{k}}(\mathbf{r})$  is then,

$$f_k(\theta) = \langle \phi_k | f(\theta, \mathbf{r}) | \phi_0 \rangle_{\mathbf{r}}$$

Achieved by replacing  $H_p \rightarrow -\epsilon_0$  in Schrödinger equation

#### Adiabatic approximation - time perspective

The time-dependent equation is

$$\mathbf{H}\Psi(\mathbf{r},\mathbf{R},\mathbf{t}) = i\hbar \frac{\partial \Psi}{\partial \mathbf{t}}$$

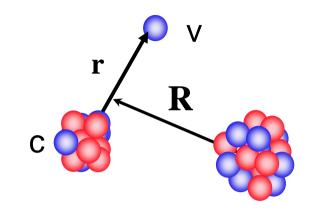
and can be written

$$\Psi(\mathbf{r}, \mathbf{R}, t) = \Lambda \Phi(\mathbf{r}(t), \mathbf{R}), \quad \mathbf{r}(t) = \Lambda^{+} \mathbf{r} \Lambda$$
  
 $\Lambda = \exp\{-i(\mathbf{H}_{p} + \epsilon_{0})t/\hbar\} \quad \text{and where}$ 

$$[T_{R} + U(\mathbf{r}(t), \mathbf{R}) - \varepsilon_{0}]\Phi(\mathbf{r}(t), \mathbf{R}) = i\hbar \frac{\partial \Phi}{\partial t}$$

Adiabatic equation

$$[T_R + U(\mathbf{r}, \mathbf{R})]\Phi(\mathbf{r}, \mathbf{R}) = (E + \varepsilon_0)\Phi(\mathbf{r}, \mathbf{R})$$



Adiabatic step assumes

 $\mathbf{r}(t) \approx \mathbf{r}(0) = \mathbf{r} = \text{fixed}$ or  $\Lambda = 1$  for the collision time  $t_{\text{coll}}$ 

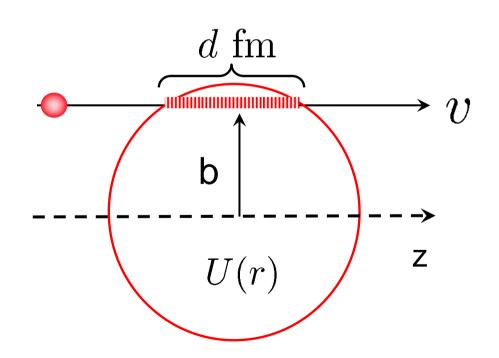
requires

$$(H_p + \varepsilon_0)t_{coll}/\hbar \ll 1$$

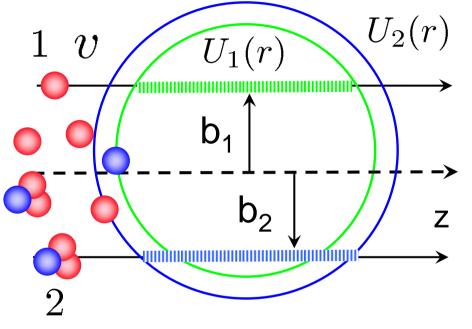
#### Reaction timescales – surface grazing collisions

For 100 and 250 MeV/u incident energy:

$$\gamma = 1.1, \ v/c = 0.42,$$
  $\gamma = 1.25, \ v/c = 0.6,$   $\Delta t = 7.9 \times d \times 10^{-24} s,$   $\Delta t = 5.6 \times d \times 10^{-24} s$ 



## Adiabatic approximation: composite projectile



$$\chi_i(b) = -\frac{1}{\hbar v} \int_{-\infty}^{\infty} U_i(r) dz$$

Total interaction energy

$$U(r_1,\ldots) = \sum_i U_i(r_i)$$

$$S_i(b_i) = \exp\left[i\chi_i(b_i)\right] = \exp\left[-\frac{i}{\hbar v} \int_{-\infty}^{\infty} U_i(r_i)dz'\right]$$

$$\chi(b_1,\ldots) = -\frac{1}{\hbar v} \int_{-\infty}^{\infty} \sum_{i} U_i(r_i) dz$$

with composite systems: get products of the S-matrices

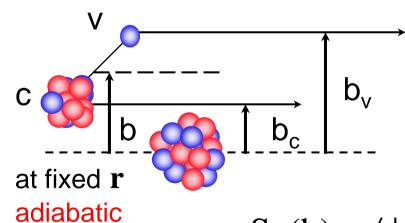
$$\exp[i\chi(b_1,\ldots)] = \prod_i S_i(b_i)$$

## Few-body eikonal model amplitudes

So, after the collision, as  $Z \rightarrow \infty$   $\omega(\mathbf{r}, \mathbf{R}) = S_{\alpha}(b_{\alpha}) S_{\nu}(b_{\nu})$ 

$$\Psi_{\mathbf{K}}^{\text{Eik}}(\mathbf{r},\mathbf{R}) \rightarrow e^{i\mathbf{K}\cdot\mathbf{R}} S_{c}(b_{c}) S_{v}(b_{v}) \phi_{0}(\mathbf{r})$$

with  $S_c$  and  $S_v$  the eikonal approximations to the S-matrices for the independent scattering of c and v from the target - the dynamics



So, elastic amplitude (S-matrix) for the scattering of the projectile at an impact parameter b - i.e. The amplitude that it emerges in state  $\phi_0(\mathbf{r})$  is

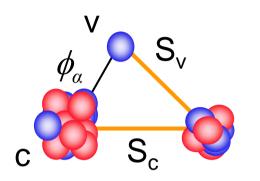
$$S_{p}(b) = \langle \phi_{0} | \underbrace{S_{c}(b_{c}) S_{v}(b_{v})}_{c} | \phi_{0} \rangle_{r}$$

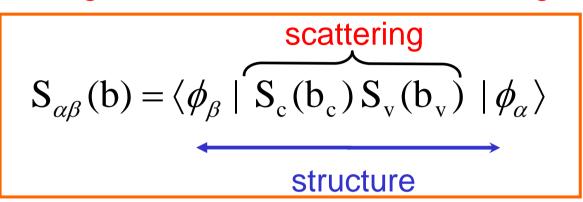
probabilities of c and v

averaged over position \_\_\_\_ amplitude that c,v survive interaction with b<sub>c</sub> and b<sub>v</sub>

#### Eikonal theory - dynamics and structure

Independent scattering information of c and v from target





Use the <u>best available</u> few- or many-body wave functions

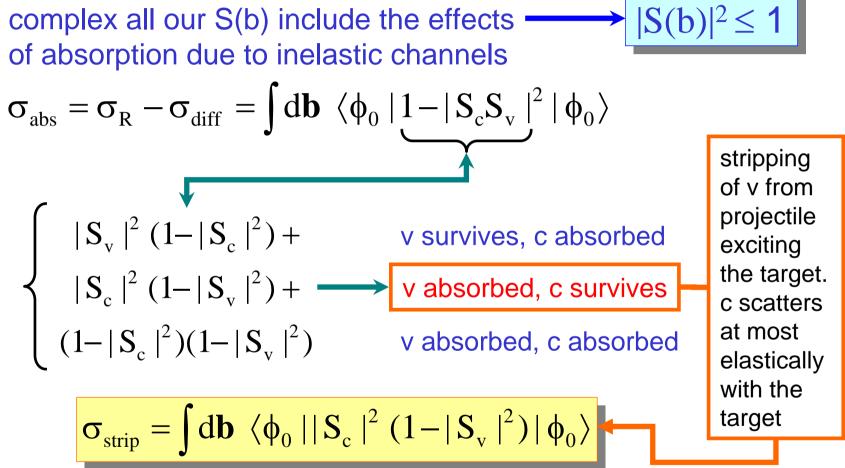
#### More generally,

$$S_{\alpha\beta}(b) = \langle \phi_{\beta} | S_1(b_1) S_2(b_2) ..... S_n(b_n) | \phi_{\alpha} \rangle$$

for any choice of 1,2 ,3, ..... n clusters for which a most realistic wave function  $\varphi$  is available

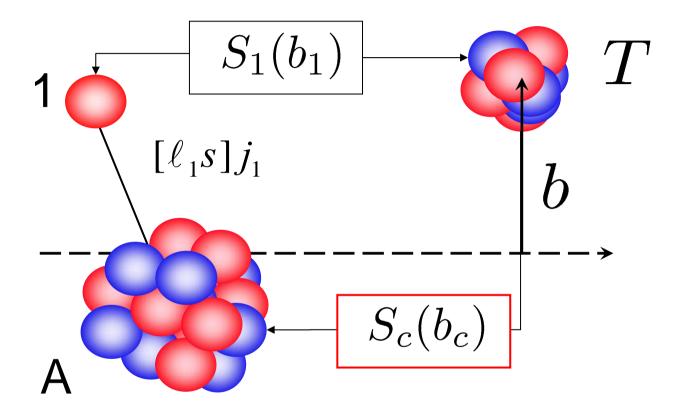
#### Absorptive cross sections - target excitation

Since our effective interactions are complex all our S(b) include the effects of absorption due to inelastic channels



Related equations exist for the differential cross sections, etc.

# Stripping of a nucleon – nucleon 'absorbed'



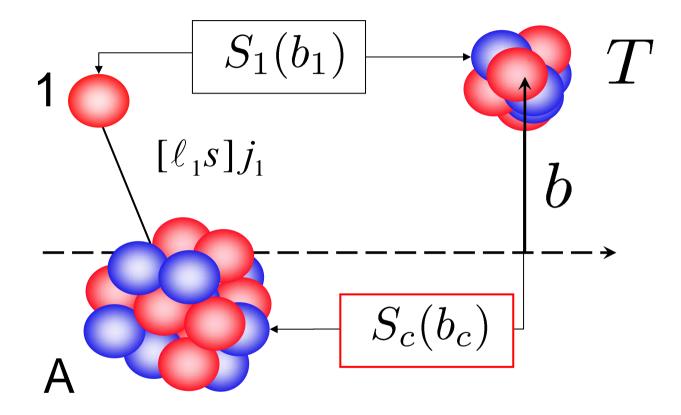
$$\sigma_{\text{strip}} = \int d\mathbf{b} \langle \phi_0 || \mathbf{S}_C |^2 (1 - |\mathbf{S}_1|^2) |\phi_0\rangle$$

#### And for two-nucleon removal

$$\sigma_{abs} \rightarrow 1 - |S_c|^2 |S_1|^2 |S_2|^2$$

$$1 = [|S_c|^2 + (1 - |S_c|^2)] 
\times [|S_1|^2 + (1 - |S_1|^2)] 
\times [|S_2|^2 + (1 - |S_2|^2)]$$
core survival
and nucleon
"removal"

#### Diffractive (breakup) removal of a nucleon



$$\sigma_{\text{diff}} = \int d\mathbf{b} \left\{ \langle \phi_0 || \mathbf{S}_c \mathbf{S}_v |^2 |\phi_0 \rangle - |\langle \phi_0 |\mathbf{S}_c \mathbf{S}_v |\phi_0 \rangle|^2 \right\}$$

# Core-target effective interactions – for S<sub>c</sub>(b<sub>c</sub>)

Double folding  $U_{AB}(R) = \int d\mathbf{r}_1 \int d\mathbf{r}_2 \ \rho_A(\mathbf{r}_1) \ \rho_B(\mathbf{r}_2) \ t_{NN}(\mathbf{R} + \mathbf{r}_2 - \mathbf{r}_1)$  R  $P_B(\mathbf{r}_2) \mathbf{r}_1 \mathbf{r}_2 \mathbf{r}_2 \mathbf{r}_3 \mathbf{r}_4 \mathbf{r}_3 \mathbf{r}_4 \mathbf{r}_4 \mathbf{r}_5 \mathbf{r}_5$ 

At higher energies – for nucleus-nucleus or nucleon-nucleus systems – first order term of multiple scattering expansion

$$t_{NN}(r) = \left[ -\frac{\hbar v}{2} \sigma_{NN}(i + \alpha_{NN}) \right] f(r), \quad \int d\vec{r} f(r) = 1$$

e.g. 
$$f(r) = \delta(r)$$

nucleon-nucleon cross section

$$f(r) = (\sqrt{\pi}t)^{-3} \exp(-r^2/t^2)$$

resulting in a COMPLEX nucleus-nucleus potential

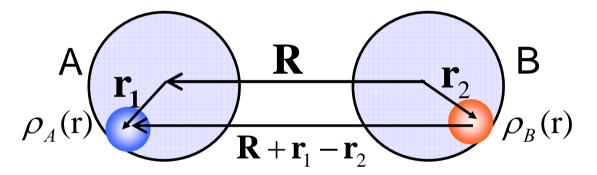
M.E. Brandan and G.R. Satchler, Phys. Rep. 285 (1997) 143-243.

# Effective interactions – Folding models

Double folding

$$\mathbf{U}_{AB}(\mathbf{R}) = \int d\mathbf{r}_1 \int d\mathbf{r}_2 \, \rho_A(\mathbf{r}_1) \, \rho_B(\mathbf{r}_2) \, \mathbf{v}_{NN}(\mathbf{R} + \mathbf{r}_1 - \mathbf{r}_2)$$

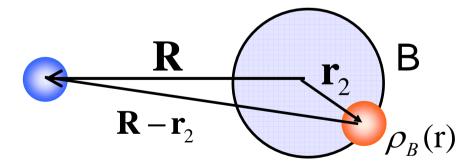
 ${
m U}_{{\scriptscriptstyle AB}}$ 



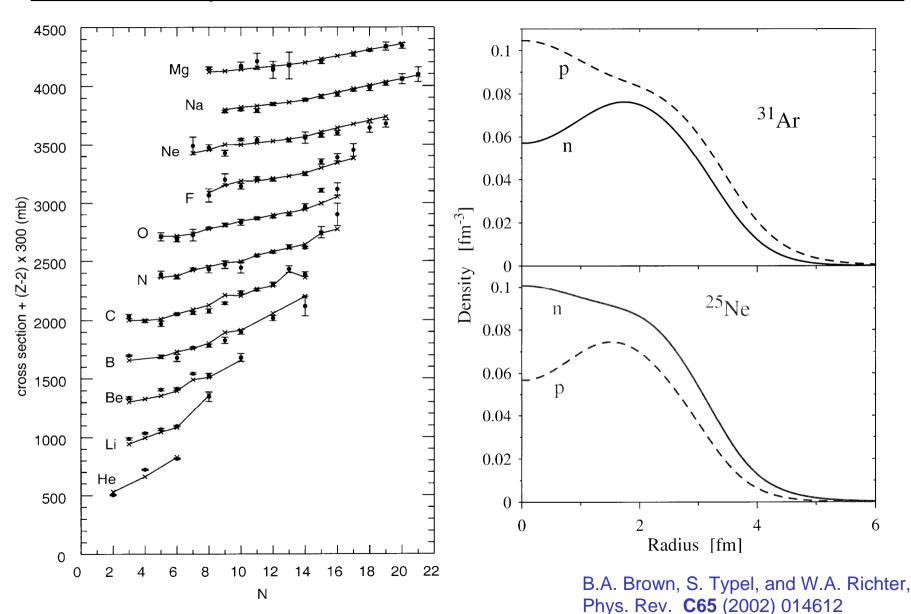
Single folding

$$\mathbf{U}_{B}(\mathbf{R}) = \int d\mathbf{r}_{2} \, \rho_{B}(\mathbf{r}_{2}) \, \mathbf{v}_{NN}(\mathbf{R} - \mathbf{r}_{2})$$

 $\mathbf{U}_{B}$ 



# Sizes - Skyrme Hartree-Fock radii and densities



# Connection to many-body – spectroscopic factors

In a potential model it is natural to define <u>normalised</u> bound state wave functions.  $A_{\mathbf{V}}(I^{\pi})$ 

$$\phi_{n\ell j}^{m}(\vec{r}) = \sum_{\lambda \sigma} (\ell \lambda s \sigma | j m) \frac{u_{n\ell j}(r)}{r} Y_{\ell}^{\lambda}(\hat{r}) \chi_{s}^{\sigma},$$

$$\int_{0}^{\infty} [u_{n\ell j}(r)]^{2} dr = 1$$

$$n\ell_{j}$$

$$A-1 \chi(J_{f}^{\pi})$$

The potential model wave function approximates the <u>overlap function</u> of the A and A–1 body wave functions (A and A–n in the case of an n-body cluster) i.e. the overlap

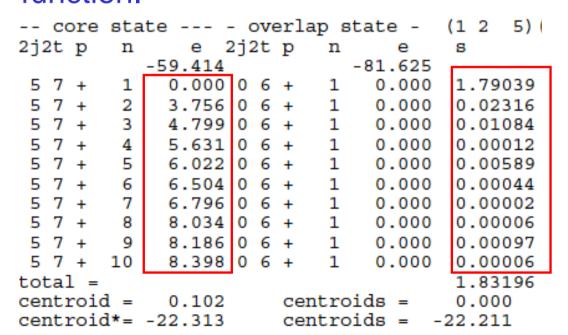
$$\langle \ell j, \vec{r}, A^{-1} \mathbf{X}(J_f^{\pi}) | A \mathbf{Y}(J_i^{\pi}) \rangle \to I_{\ell j}(r), \quad \int_0^{\infty} [I_{\ell j}(r)]^2 dr = S(J_i, J_f \ell j)$$

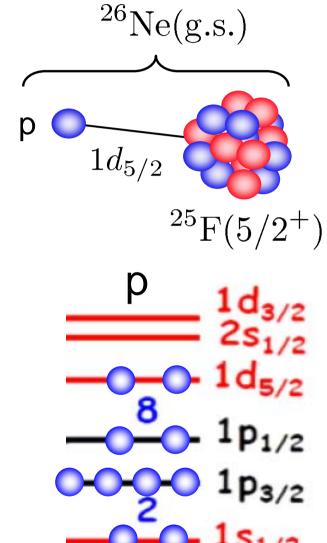
S(...) is the <u>spectroscopic factor</u>  $\leftarrow$  a structure calculation

## Many-body input – the shell model overlaps

$$\langle \vec{r},^{25} \text{Ne}(5/2^+, E^*)|^{26} \text{Ne}(0^+, \text{g.s.}) \rangle$$

USDA sd-shell model overlap from e.g. OXBASH (*Alex Brown et al.*). Provides spectroscopic factors but not the bound state radial wave function.





We have the background we need to apply this in a serious context and confront experiment: tomorrow we look at a number of applications.

